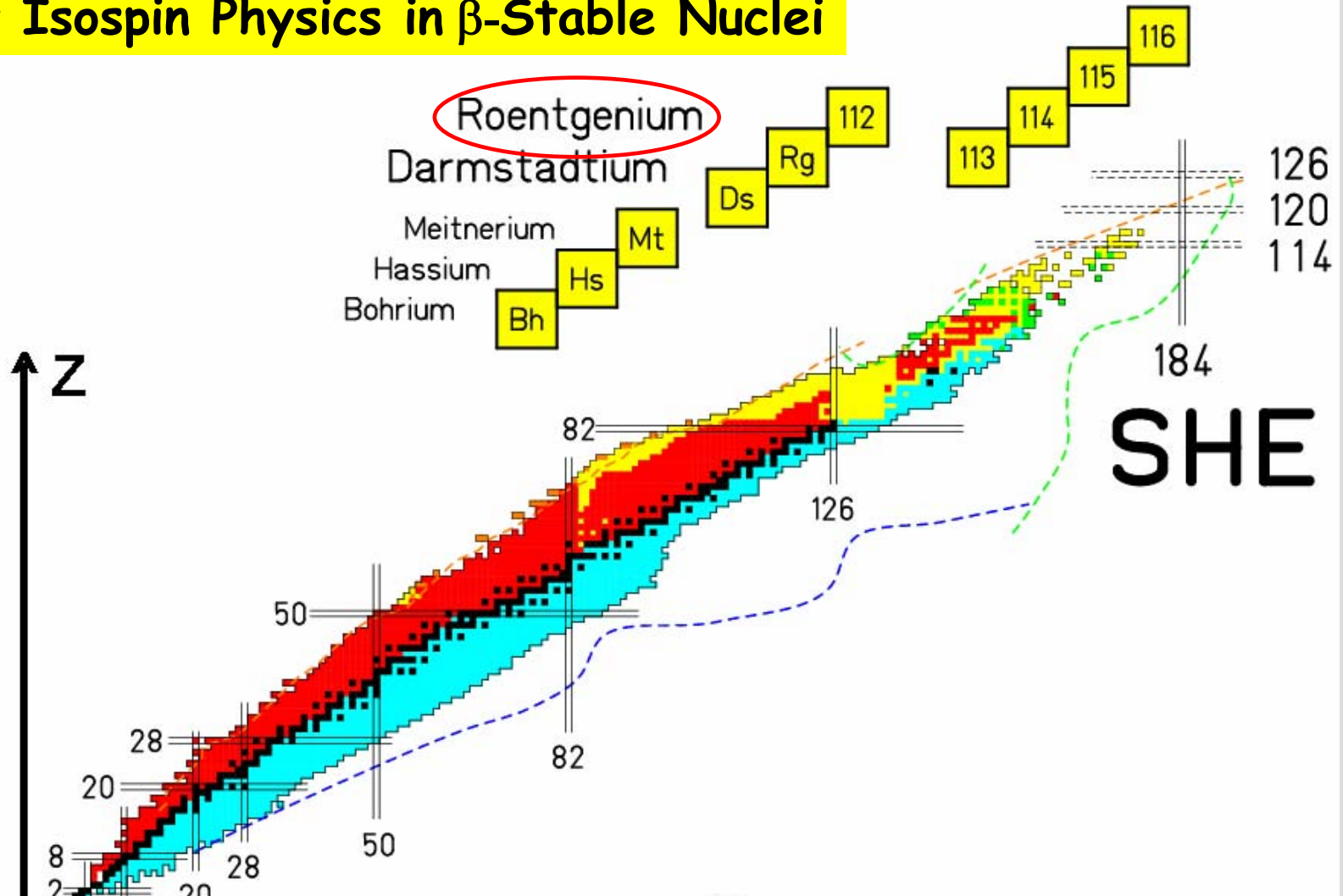


The first Century of Nuclear Physics: Low Isospin Physics in β -Stable Nuclei



$$E(A)/A = -16\text{MeV} + E_{\text{surf}}/A^{1/3} + E_{\text{coul}}$$

$$+ [(N-Z)/A]^2(a_4 + C_{\text{sym}}/A^{1/3}) + E_{\text{pair}} + E_{\text{shell}}$$

The second Century: Physics of Asymmetric Nuclear Matter

H. Lenske
Institut für Theoretische Physik, U. Giessen



EuroGK



SFB/TR 16

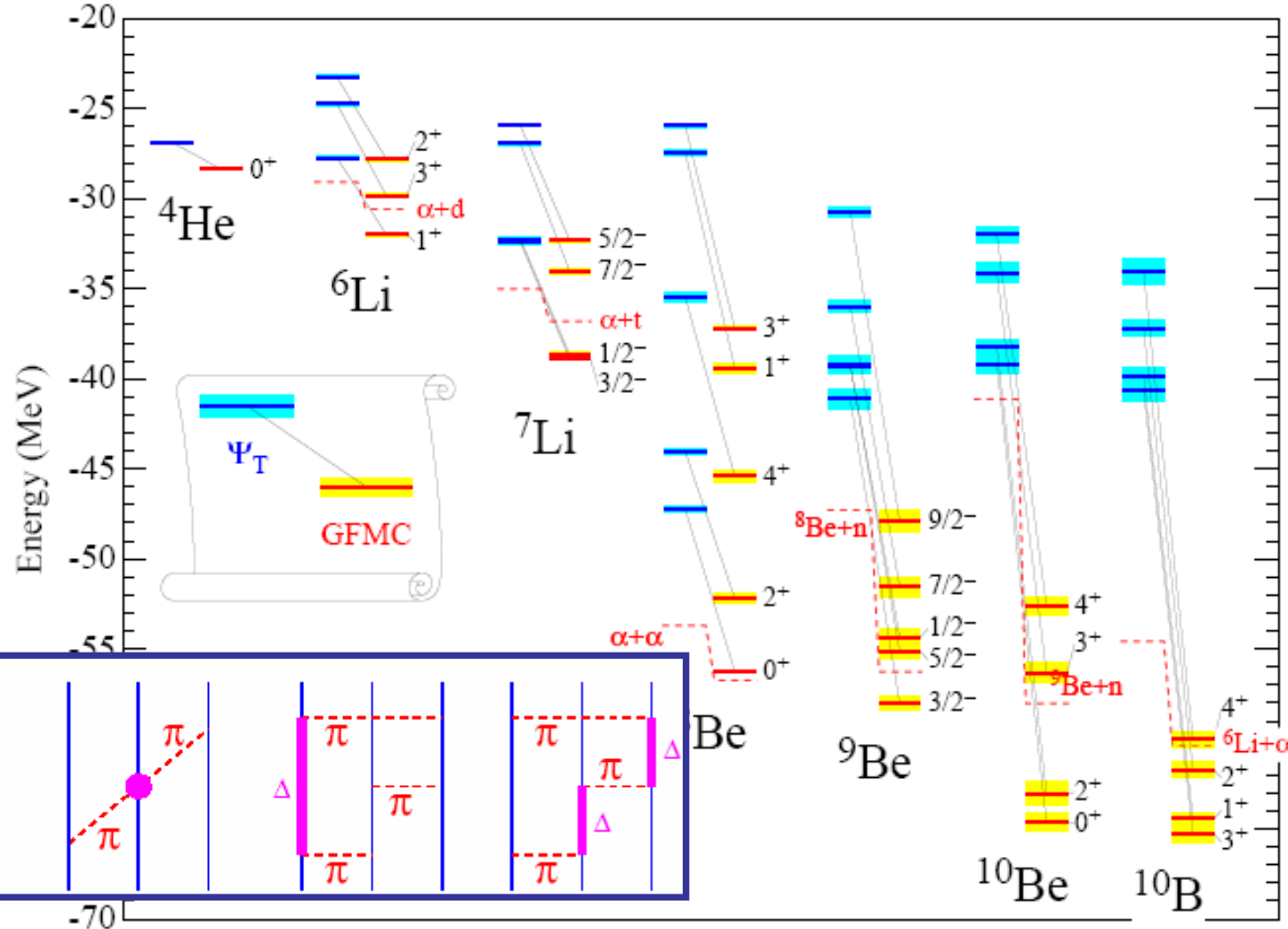
The Agenda:

- Our position : Competing with LHC physics!
- Our goal : Define intellectually demanding Tasks
- Our business : The Physics of Nuclear Matter
- Our wish : Nucleo-Engineering on the Femto-scale

The Topics

- The Fate of Mean-Field Dynamics and Shell Structures?
- Scanning the Isospin Degree of Freedom
- *ab initio* Approaches to Nuclear Structure and Reactions
- Pairing and other correlations in Exotic Nuclei
- New features: Continuum Coupling and Open Quantum Systems
- The End: Summary and Outlook

Modern Nuclear Theory: *ab initio* Calculations



A=4-10 GFMC results: AV18 NN potential + IL2 Urbana 3-Body Interactions

Kohn-Sham-Theorem & Density Functional for Nuclei:

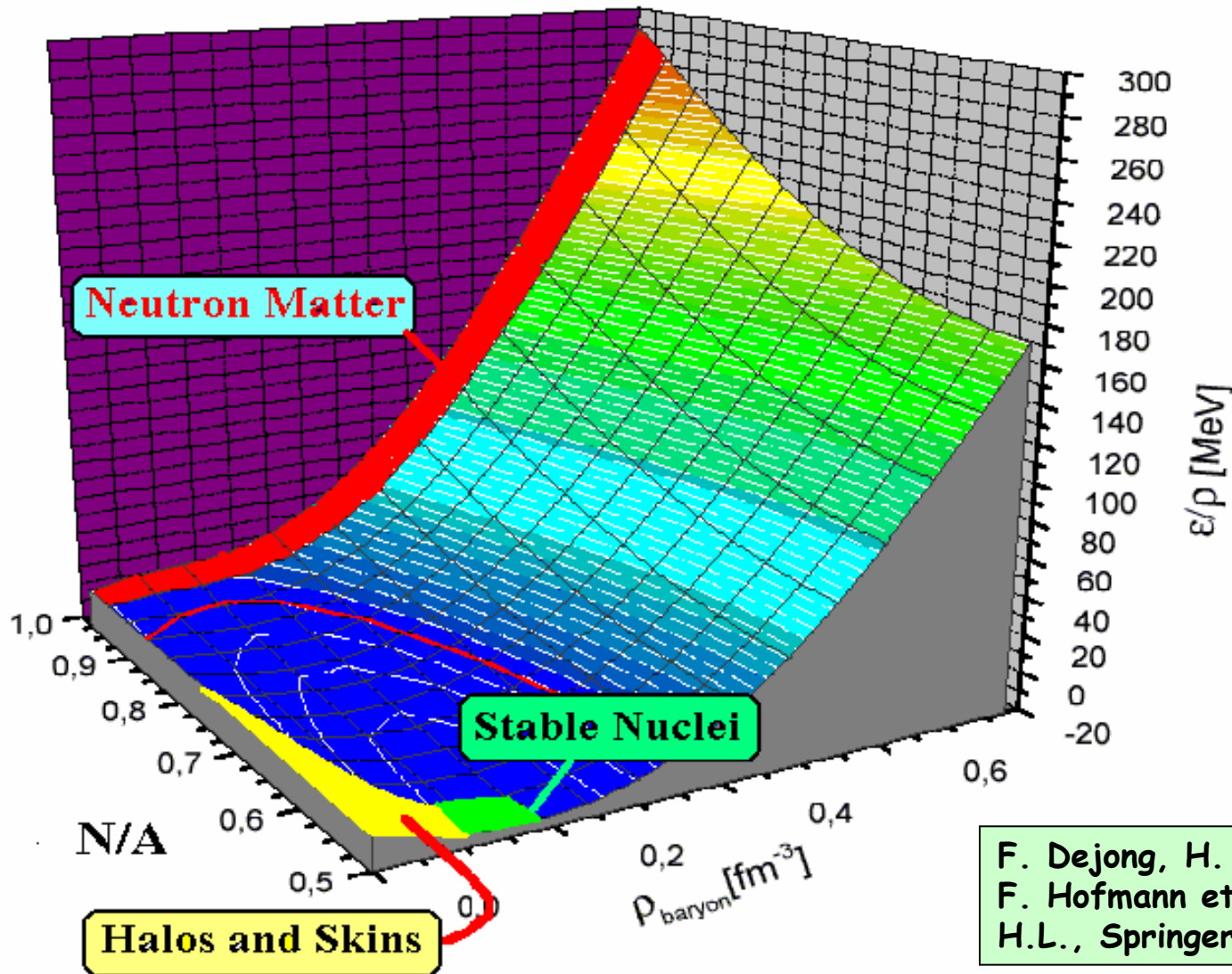
- Hohenberg-Kohn: DFT $\sim E[\rho]$
- Kohn-Sham : DFT $\sim E[\rho, \tau]$
- Nuclei : DFT $\sim E[\rho_p, \rho_n, \tau_p, \tau_n, \dots]$

$$E[\rho, \tau] = \rho\tau + \frac{1}{2}\rho^2 \frac{3}{16} \left(a_{SE}(\rho) + a_{TE}(\rho) \right) \frac{4\pi\hbar^2}{M^*(\rho)} + \dots$$

At $\rho=0$: $a_{SE}(S=0, l=1) = -23.8\text{fm}$; $a_{TE}(S=1, l=0) = +5.42\text{fm}$

$$\tau = \tau_p + \tau_n; \quad \rho = \rho_p + \rho_n$$

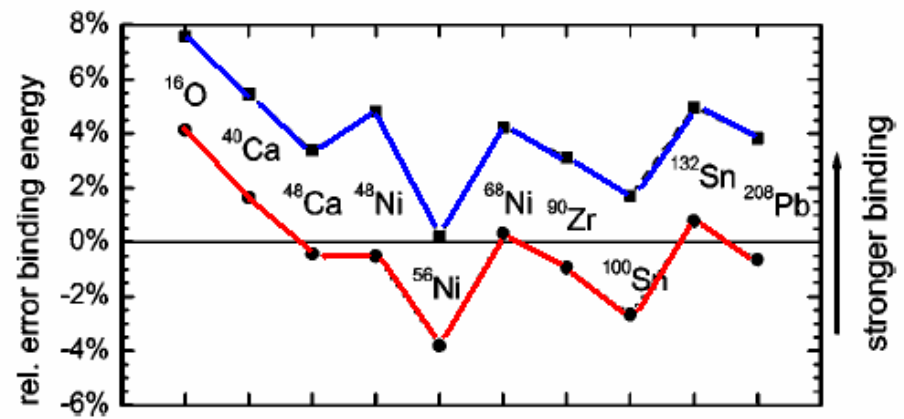
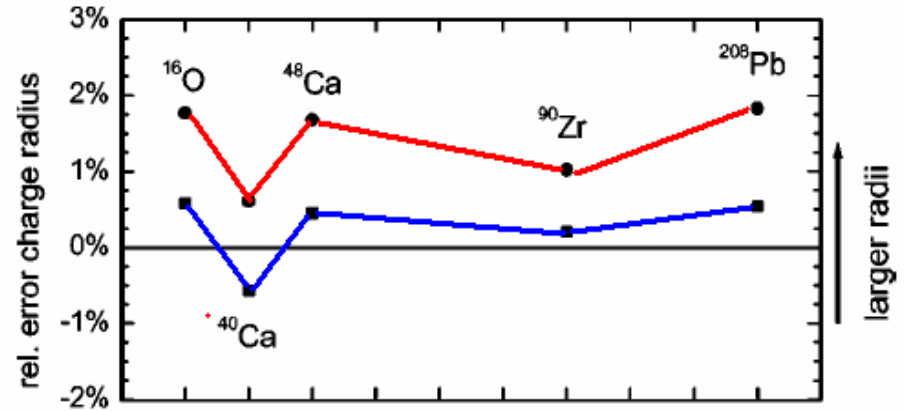
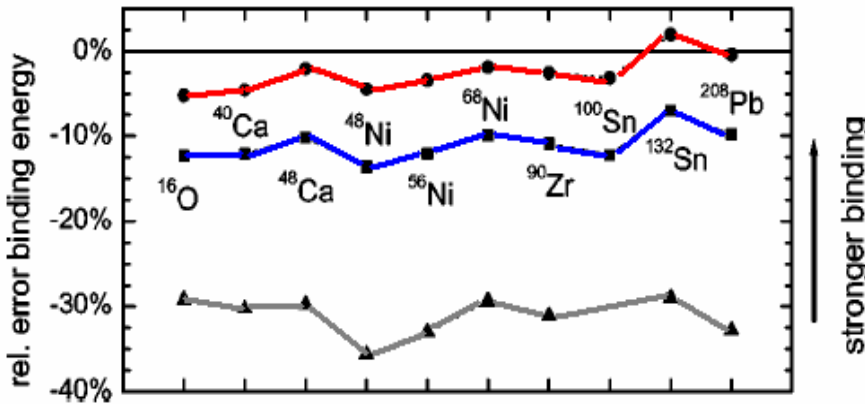
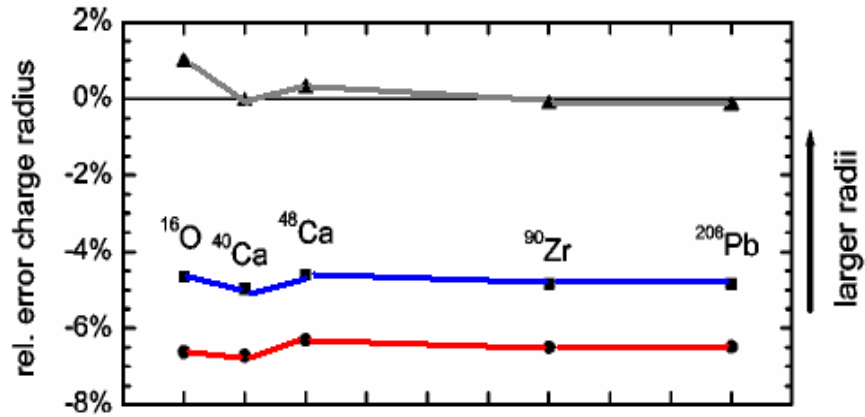
EoS of Asymmetric Nuclear Matter



DDRH
Relativistic
DFT
(Groningen NN-
Potential)

F. Dejong, H. L. PRC58 (1998),
F. Hofmann et al. PRC64 (2001)
H.L., Springer Lecture Notes 2004

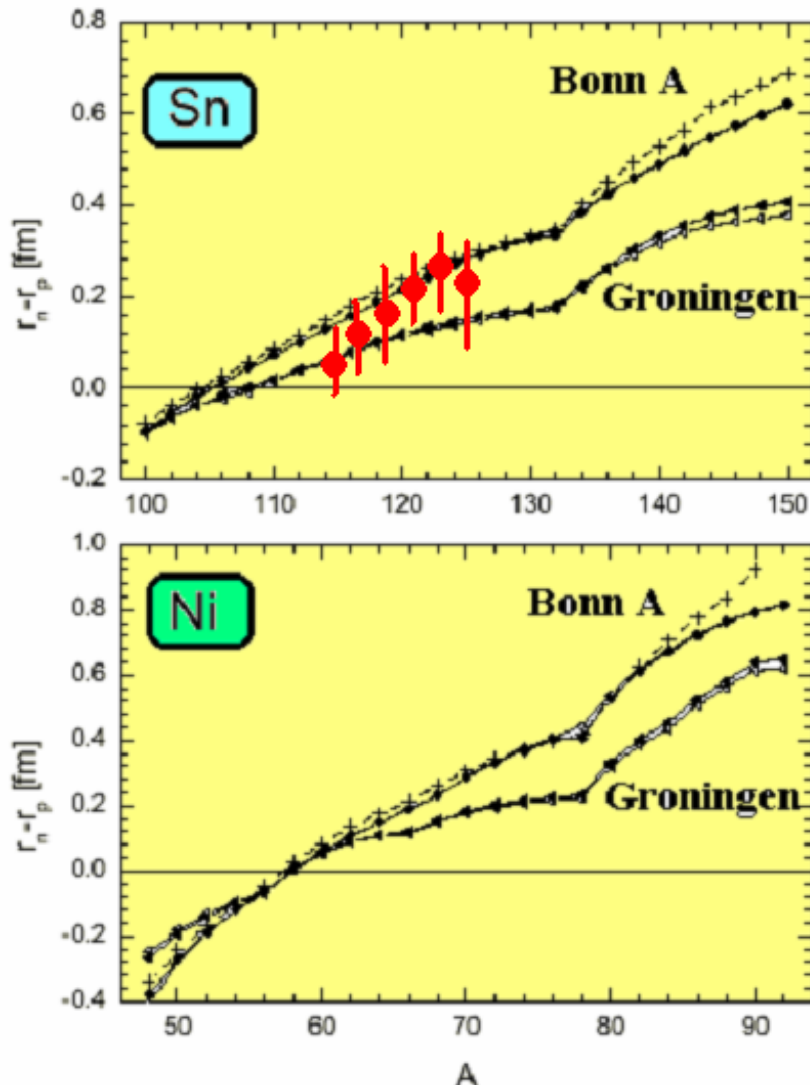
DDRH Results for Binding Energies and Charge Radii



Groningen NN Potential

Bonn A NN Potential

Neutron Skins in Ni and Sn Isotopes



DDRH RMF-Calculations
Dirac-Brueckner In-Medium Vertices
Bonn-A and Groningen NN-Potentials

Neutron Skin and Symmetry Energy:

Bonn A : $a_4 = 32$ MeV

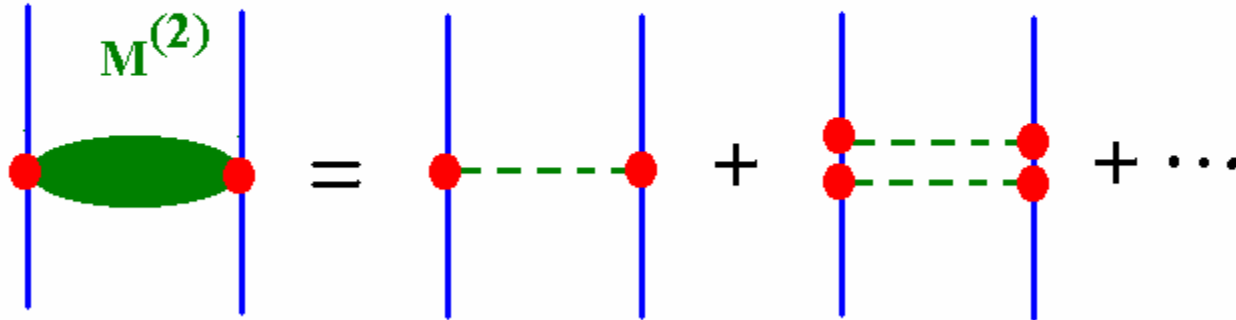
Groningen : $a_4 = 26$ MeV

Sn Data: Krasnahorkay et al. PRL 82 (1999) 3216
(from Charge Exchange Spin-Dipole sum rules)

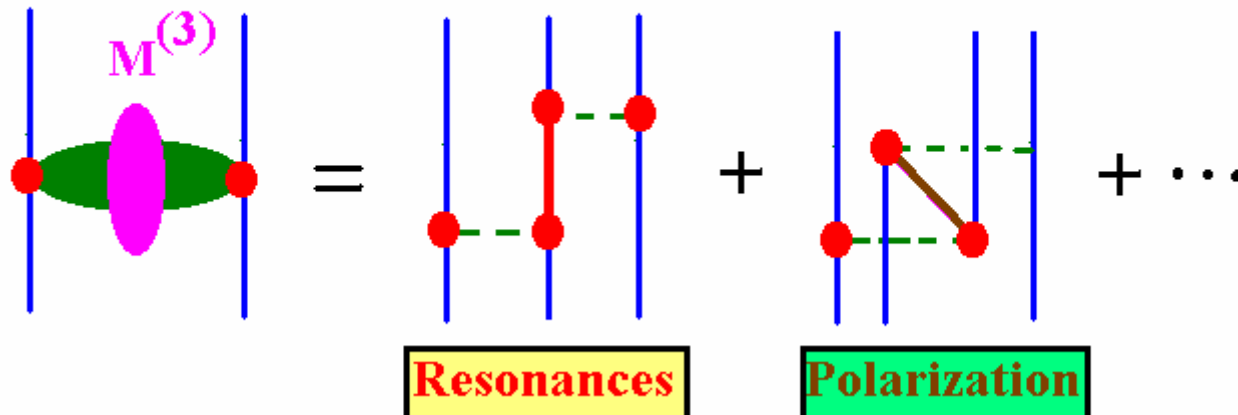
F. Hofmann et al., PR C64 (2001)
N. Tsoneva. H.L., PLB586 (2004)
N. Tsoneva. H.L., PRC (2008) in print

In-Medium Interactions

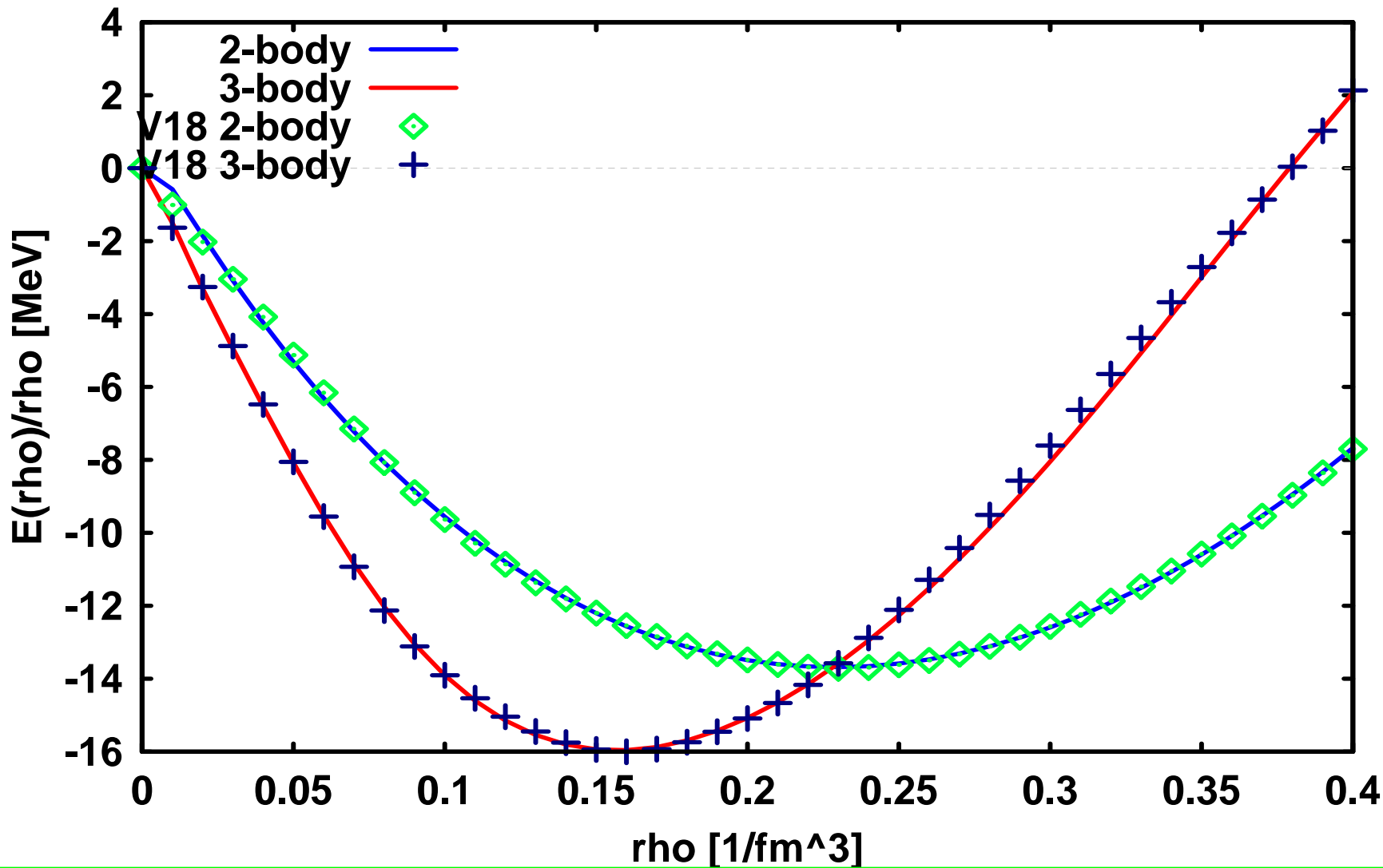
2-body Ladder (Brueckner) :



3-body interactions:

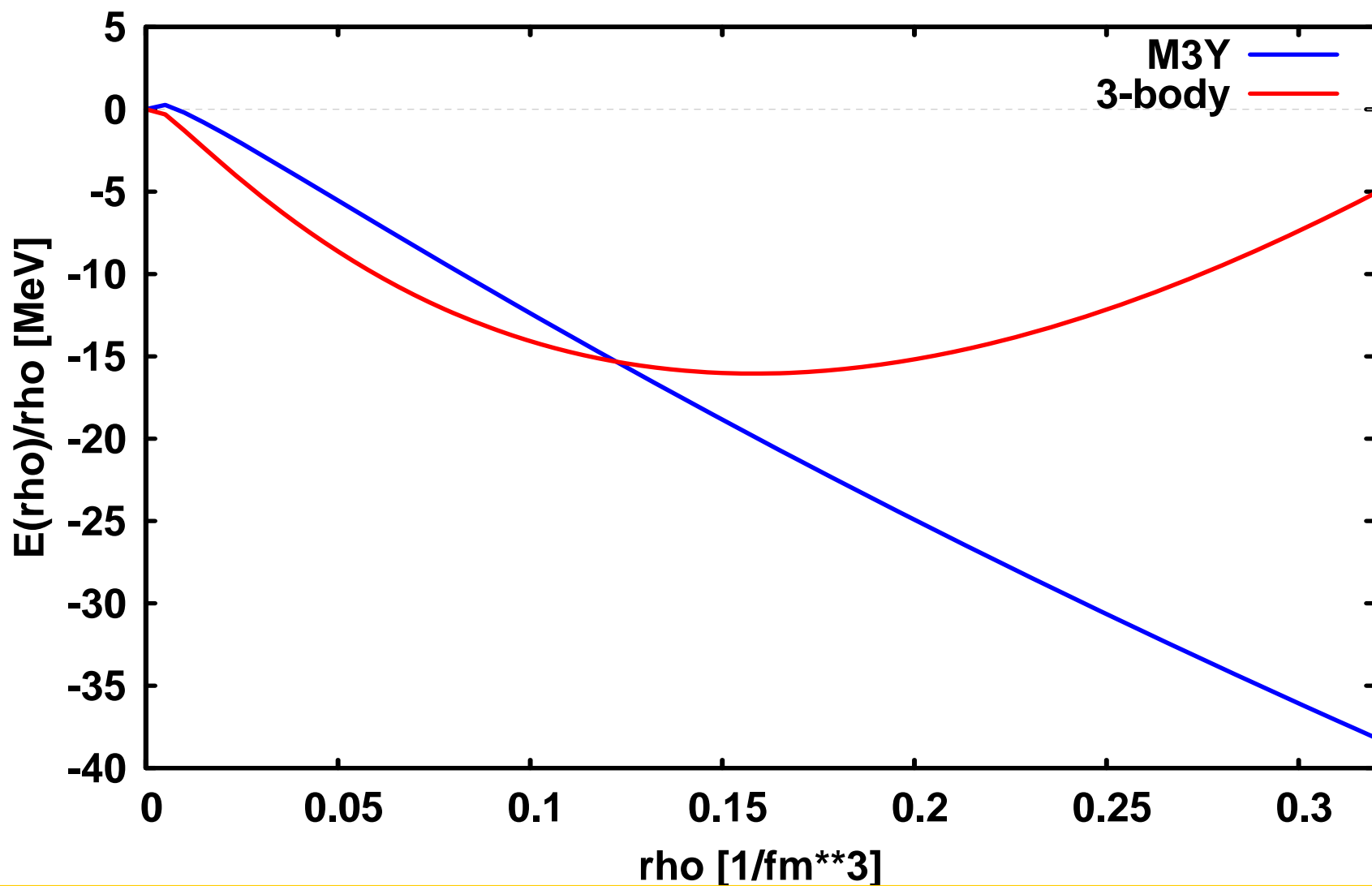


EoS of Symmetric Matter



A. Ataie, H.L., in preparation

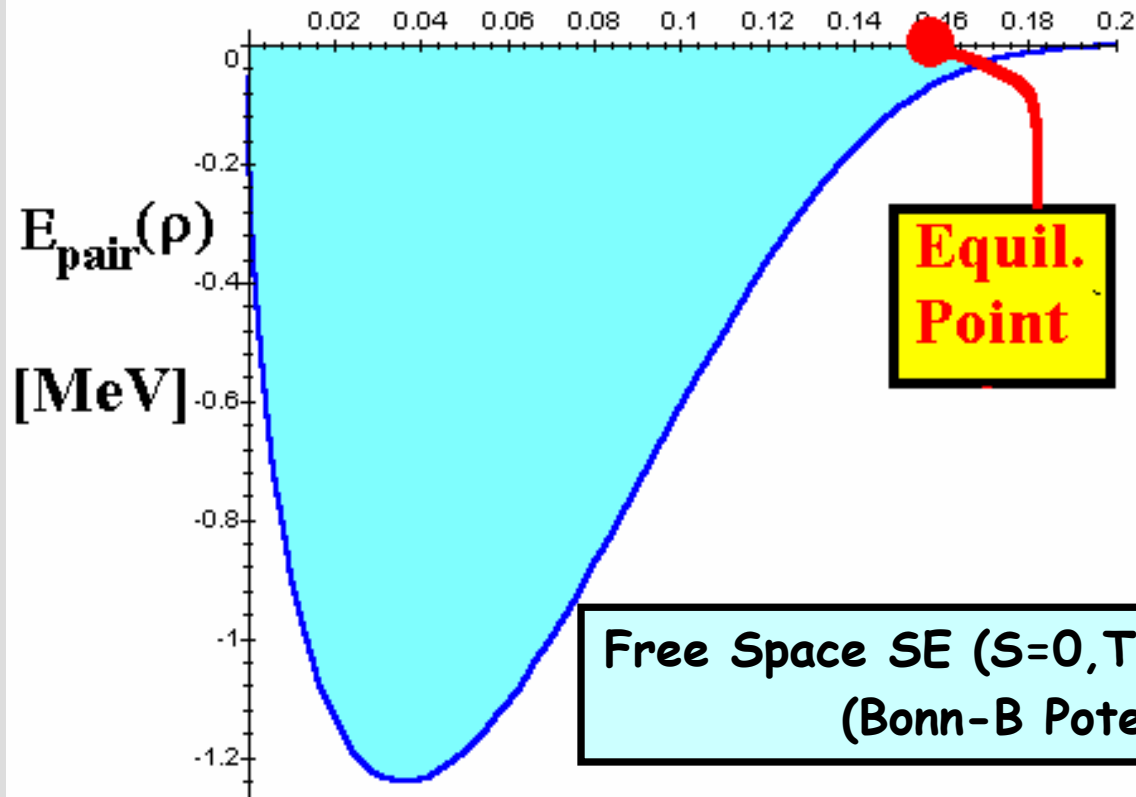
EoS from original M3Y-Parameterization



Pairing in Infinite Nuclear Matter

Pairing Energy per Particle

ρ [1/fm³]



**Equil.
Point**

Free Space SE ($S=0, T=1$) Interaction:
(Bonn-B Potential)

Pairing is a **LOW DENSITY** Phenomenon

Extended HFB Theory as Coupled Channels Problem: The Gorkov-Equations

$$\begin{pmatrix} H - \lambda & -\Delta \\ -\Delta^+ & -(H - \lambda) \end{pmatrix} \begin{pmatrix} \phi_+ \\ \phi_- \end{pmatrix} = E \begin{pmatrix} \phi_+ \\ \phi_- \end{pmatrix}$$

$$\phi_+ \sim u_{lj}^{(q)}(r) |(\ell s) jm\rangle; \quad \phi_- \sim v_{lj}^{(q)}(r) |(\ell s) jm\rangle$$

Mean - Field Hamiltonian (q = p, n):

$$H = -\frac{\hbar^2}{2m} \nabla^2 + U(\rho)$$

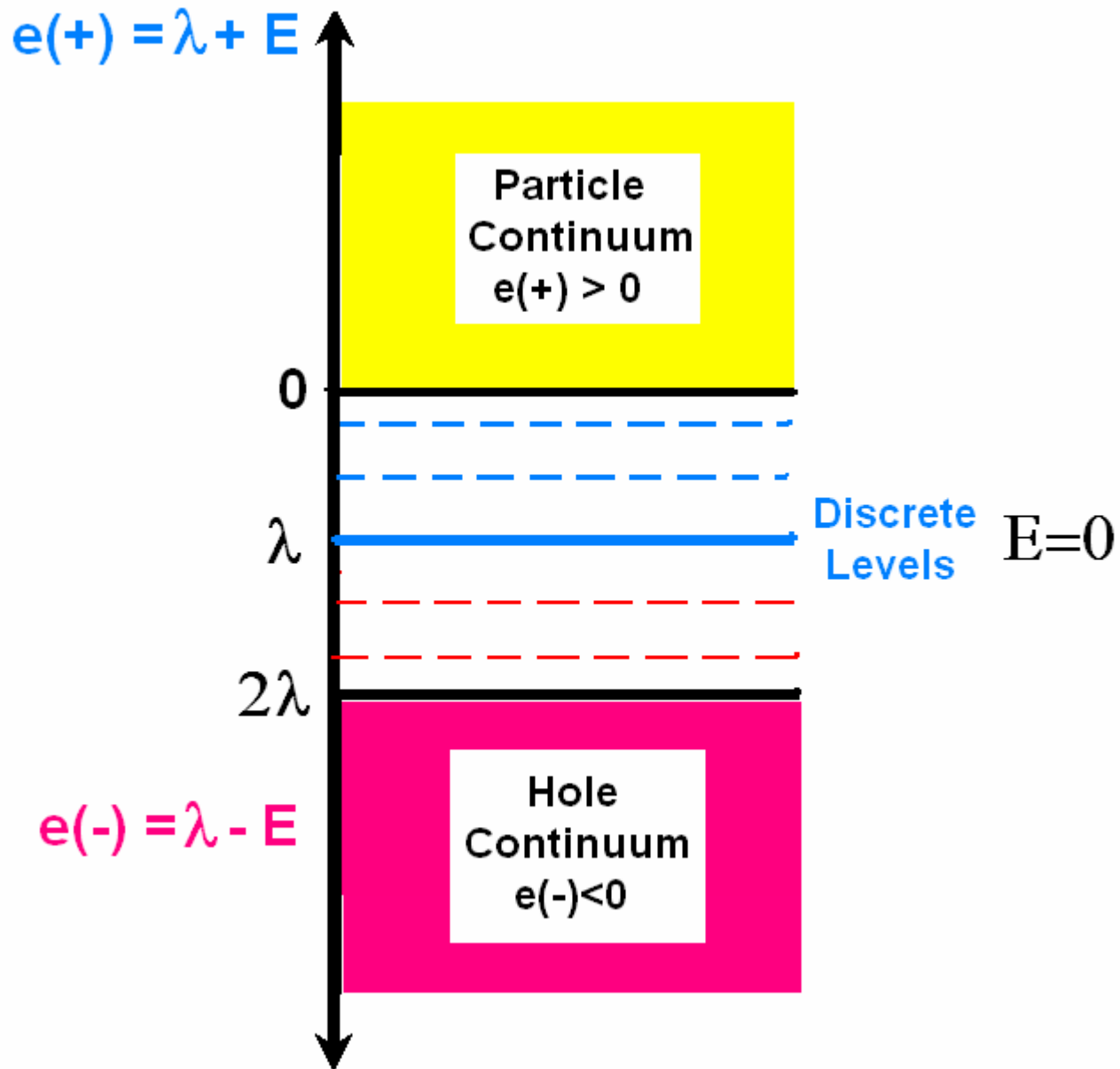
$$\rho_q(r) = \sum_{nlj} \frac{2j+1}{4\pi} |v_{nlj}^{(q)}(r)|^2$$

Pairing-Field & Density (q = p, n):

$$\Delta_q = \frac{1}{2} V_{SE} \kappa_q$$

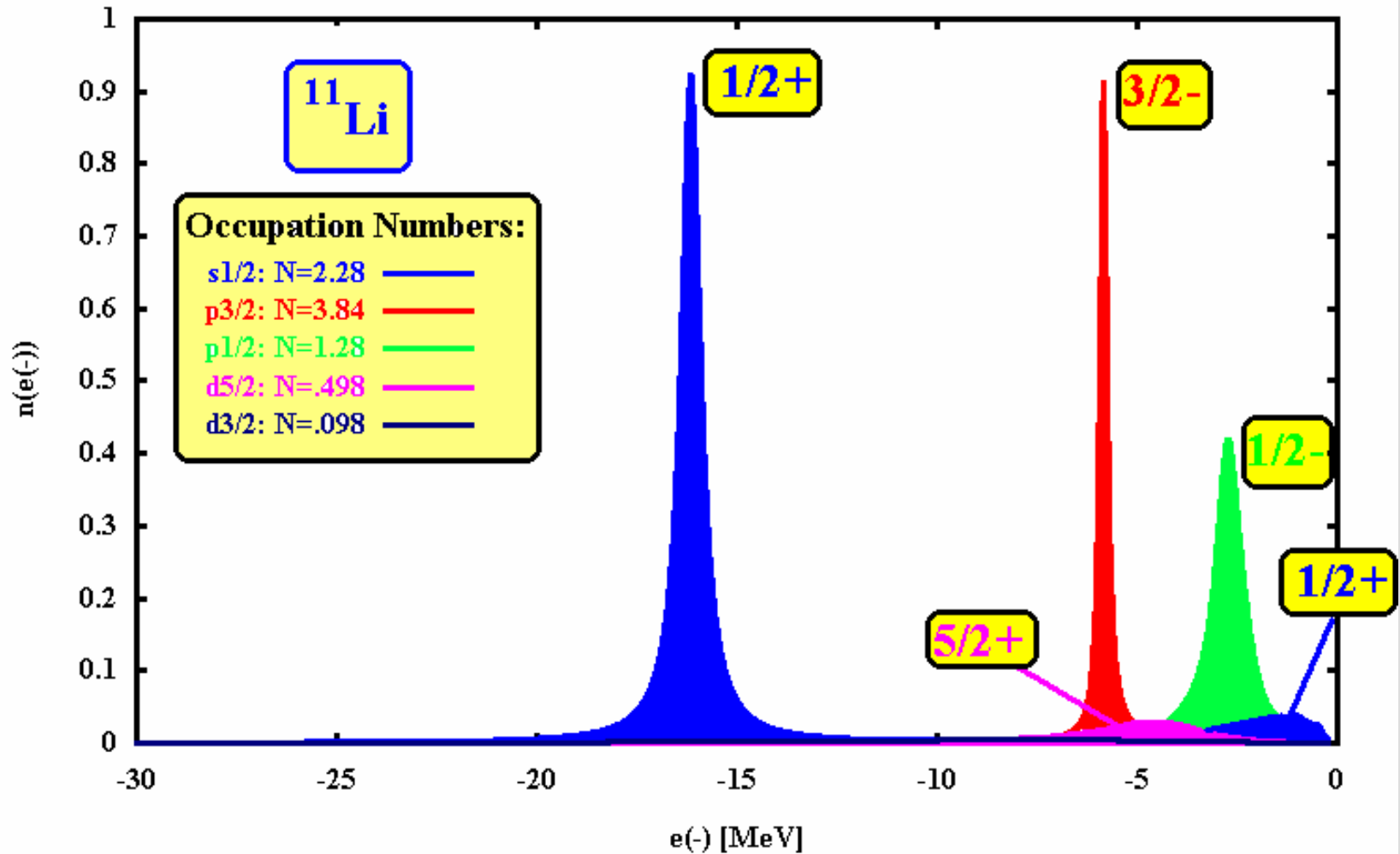
$$\kappa_q(r) = \sum_{nlj} \frac{2j+1}{4\pi} u_{nlj}^{(q)}(r) v_{nlj}^{(q)*}(r)$$

Spectrum of the Gorkov Equation:

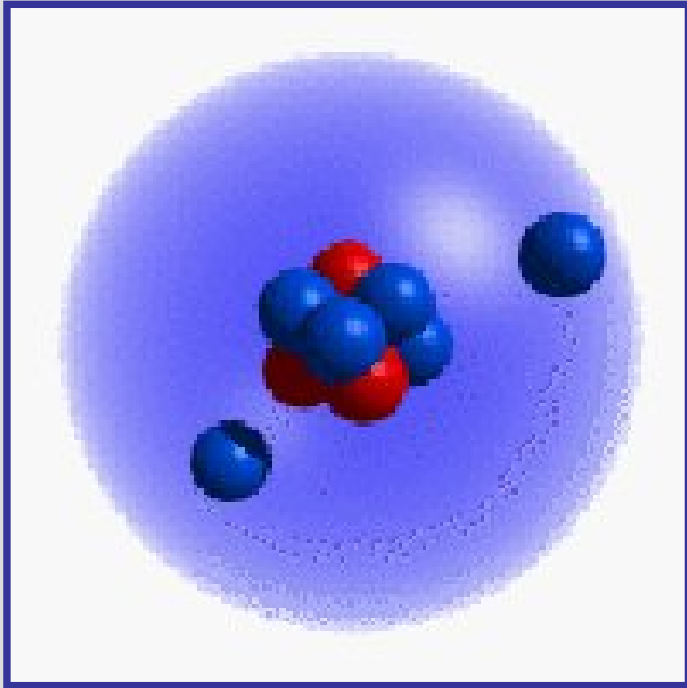


^{11}Li : Continuum HFB Spectral Functions

Neutron Spectrum :

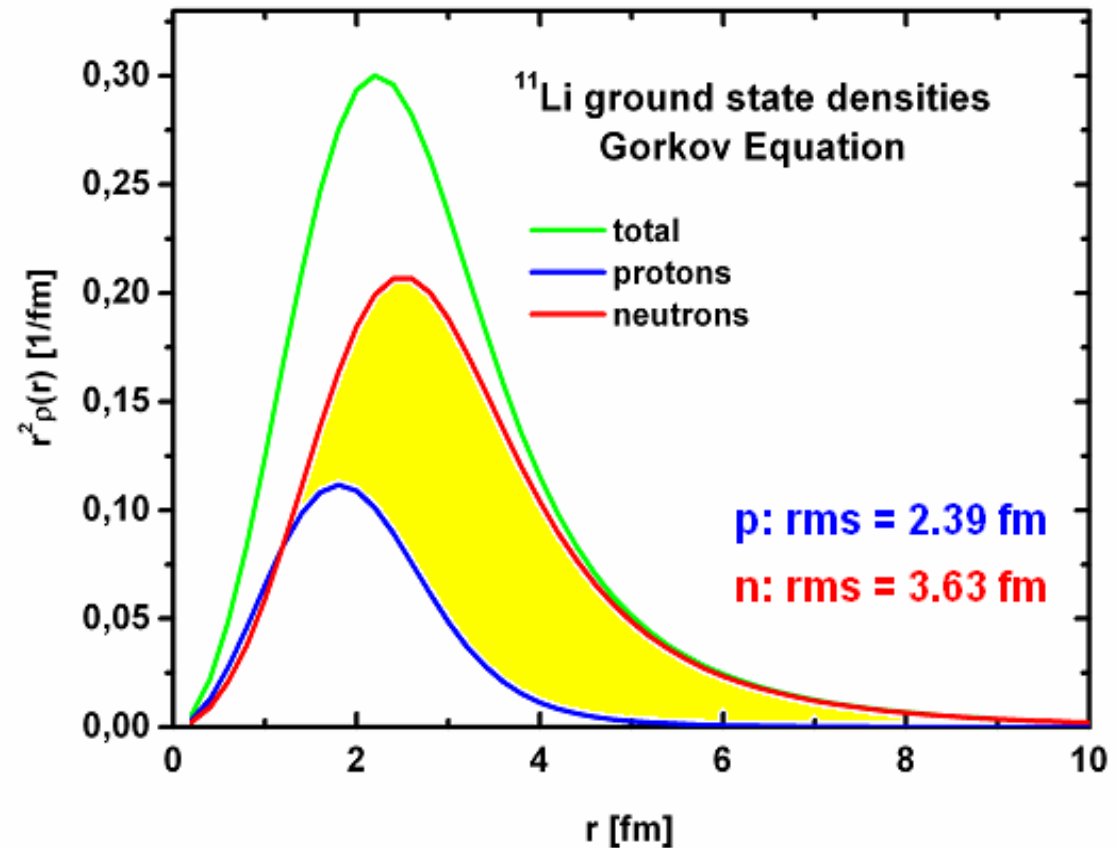


Nuclear Structure far off Stability: 2n-Halo in ^{11}Li



A popular but misleading classical point of view

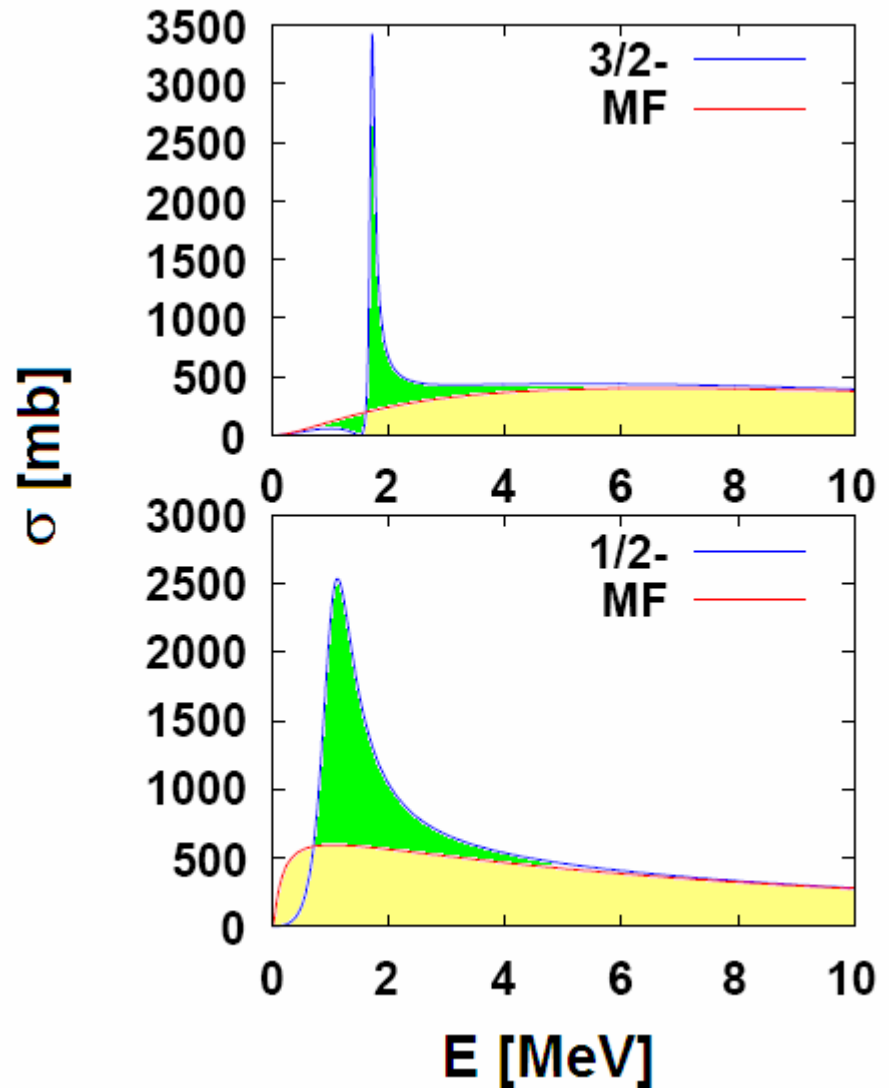
Halos: pure **Quantum Mechanics**



Pairing in the Continuum:

- CC Effects \rightarrow Interference
- Sharp resonances close to threshold
- Interaction of closed and open channels
- Admixtures of continuum strength into the ground state

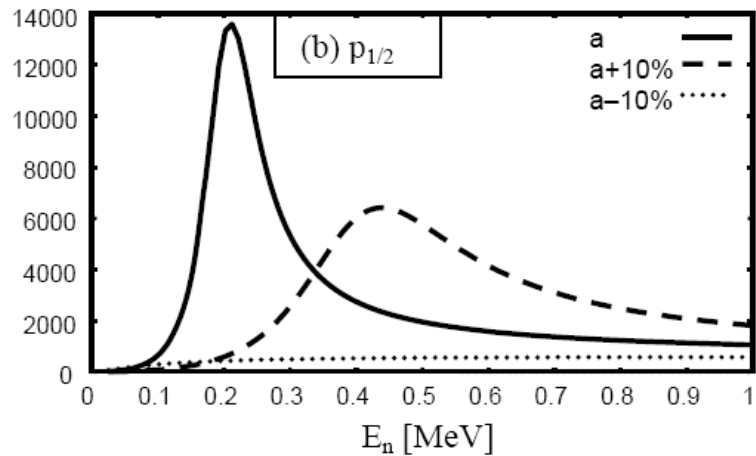
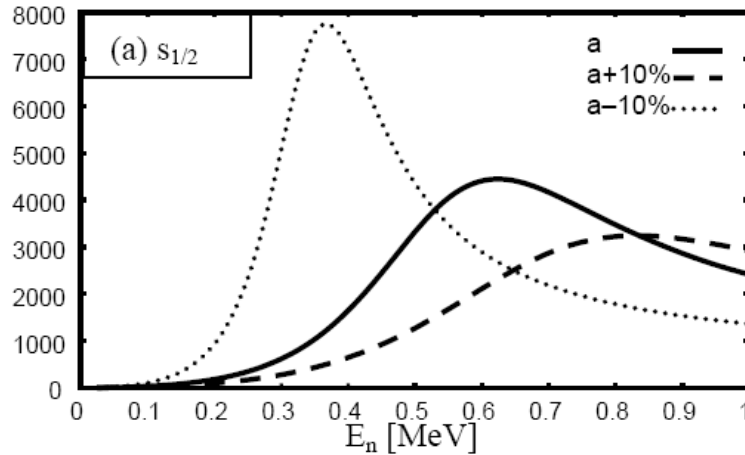
Sonja Orrigo, H.L., in preparation



Transfer to the Continuum: ${}^9\text{Li}+d \rightarrow {}^{10}\text{Li} + p$

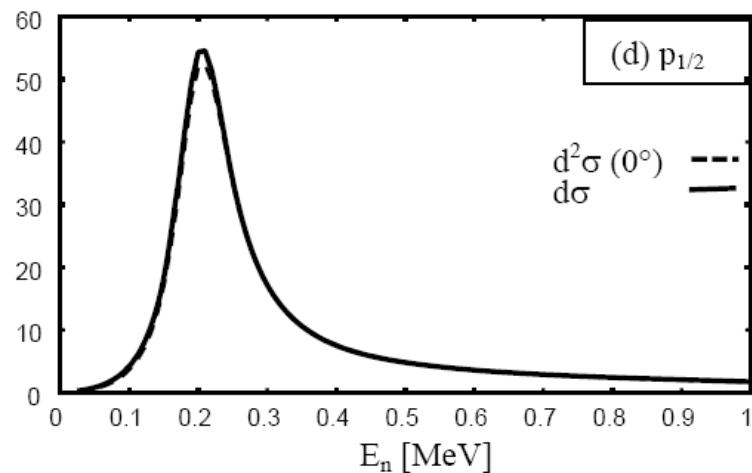
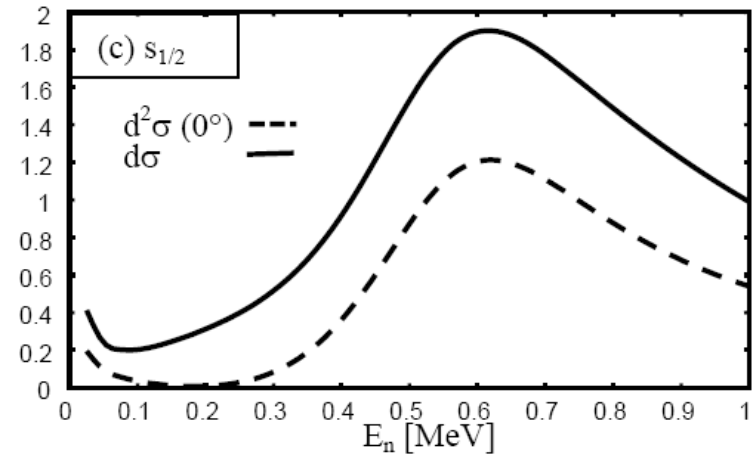
Elastic cross section

$$\frac{d\sigma_e}{dE}(E_n) \left[\frac{\text{mb}}{\text{MeV}} \right]$$



Reaction cross section

$$\frac{d^2\sigma_r}{d\Omega dE}(\theta = 0^\circ, E_n) \left[\frac{\text{mb}}{\text{sr} \cdot \text{MeV}} \right] \text{ and } \frac{d\sigma_r}{dE}(E_n) \left[\frac{\text{mb}}{\text{MeV}} \right]$$



Beyond the Ground State:

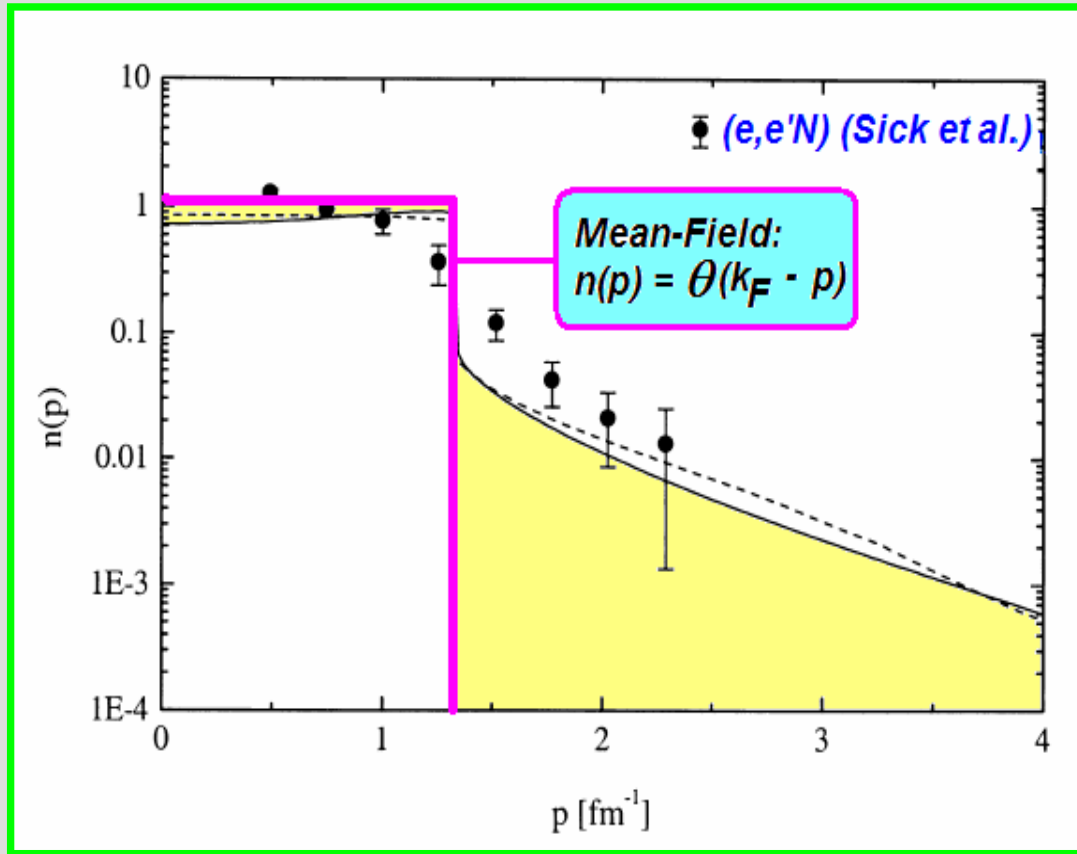
Adding Dynamics, Testing the Restoring Forces

$$E(\rho) \approx E(\rho_0) + \sum_{q=p,n} \left. \frac{\partial E(\rho)}{\partial \rho_q} \right|_{\rho_0} \delta \rho_q + \sum_{q,q'=p,n} \left. \frac{\partial^2 E(\rho)}{\partial \rho_q \partial \rho_{q'}} \right|_{\rho_0} \delta \rho_q \delta \rho_{q'} + \dots$$

$$E(\rho) \approx E(\rho_0) + \sum_{q=p,n} U_q(\rho_0) \delta \rho_q + \sum_{q,q'=p,n} F_{qq'}(\rho_0) \delta \rho_q \delta \rho_{q'} + \dots$$

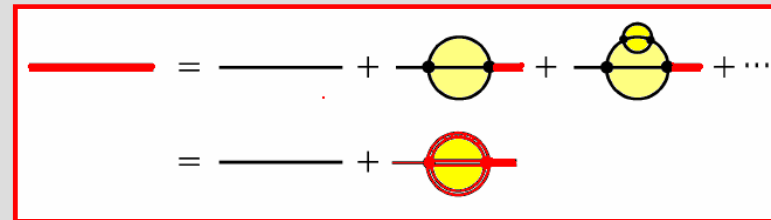
Fermi-Liquid Theory

Beyond the Mean-Field: Short-range Correlations in Nuclear Matter



Momentum Distribution

$$n(p) = N(k_F) \int \mathbf{a}(p, \omega) d\omega$$



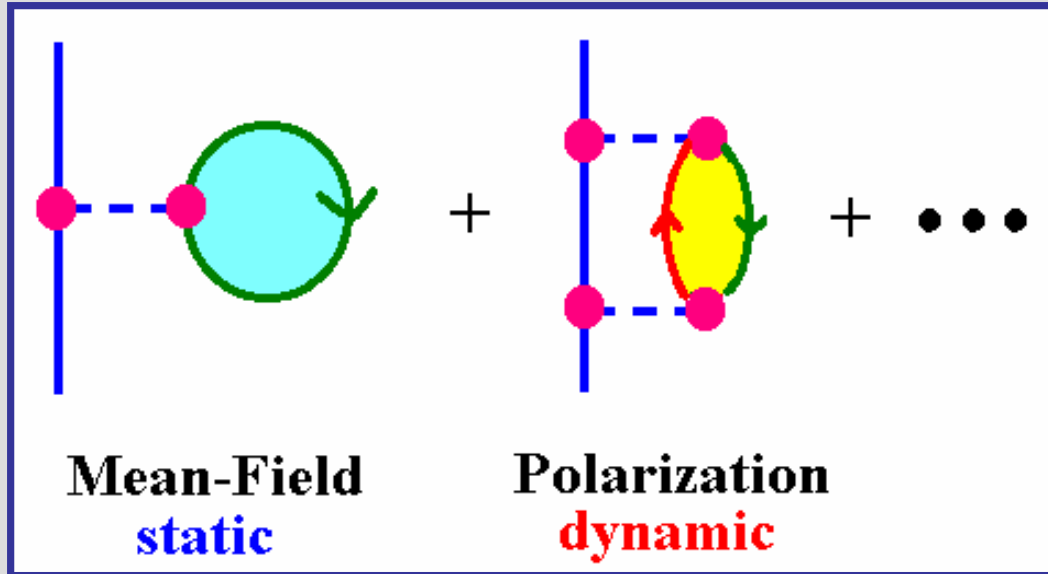
Dynamical Self - Energies :

$$g(\omega, q) = \frac{1}{\epsilon_{MF}(q) + \Sigma^{\text{ret}}(\omega, q) - \omega \pm i\eta}$$

$$\Sigma^{\text{ret}}(\omega, q) = U(\omega, q) - i \frac{1}{2} \Gamma(\omega, q)$$

$$\Gamma(\omega, q) \rightarrow 0 \text{ for } \omega \rightarrow \omega_F$$

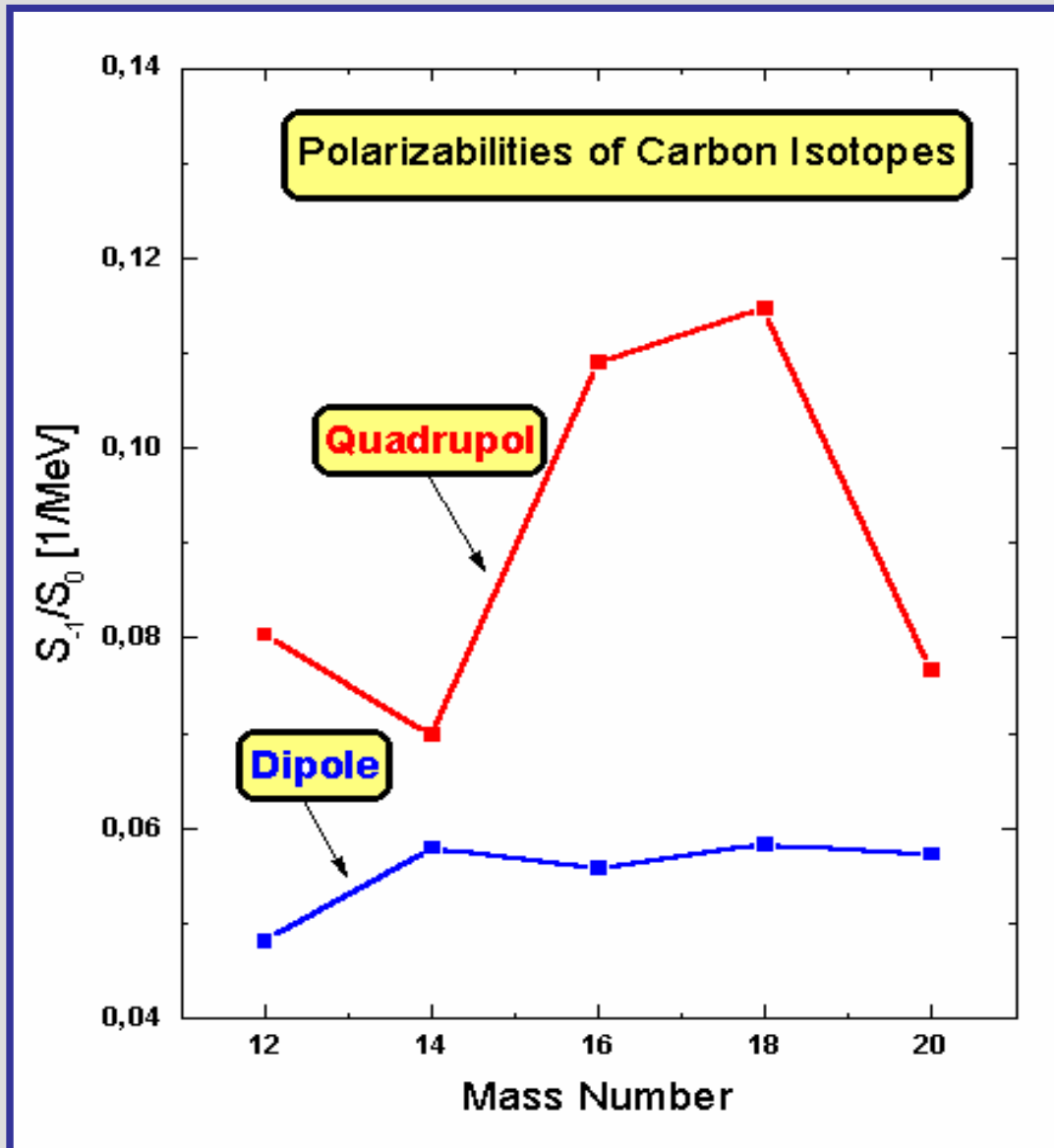
Beyond Mean-Field Dynamics: Dynamical Correlations



$$| \beta j_\beta \rangle = \sum_n z_n^\beta | n j_\beta \rangle + \sum_{kc} z_{kc}^\beta | (kc) j_\beta \rangle$$

- **Mixing** of Mean-Field s.p. States

The Softness of Exotic Nuclei



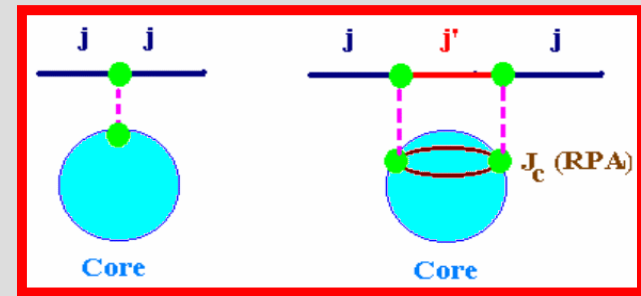
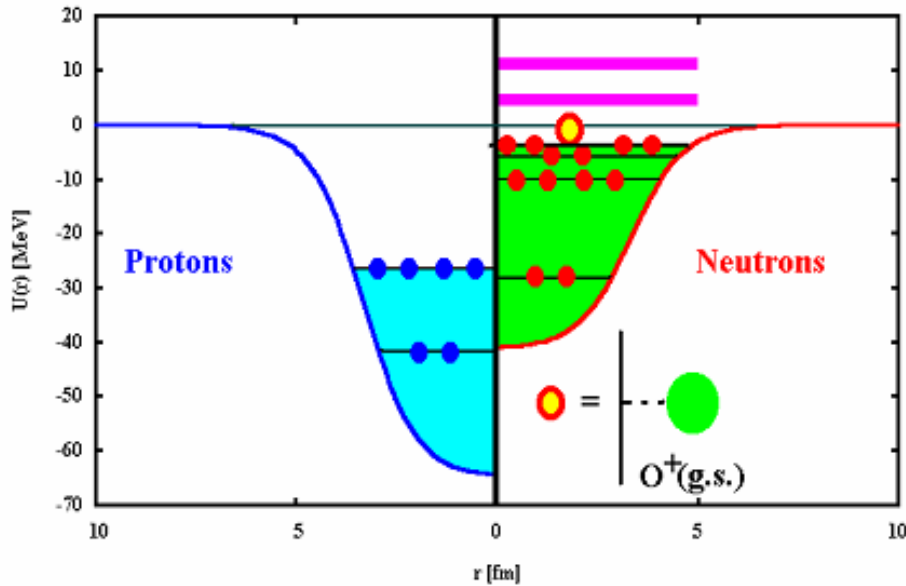
HFB & QRPA
Calculations

Polarizabilities
from Sum Rules:

$$P_{\lambda} = S_{-1}(\lambda)/S_0(\lambda)$$

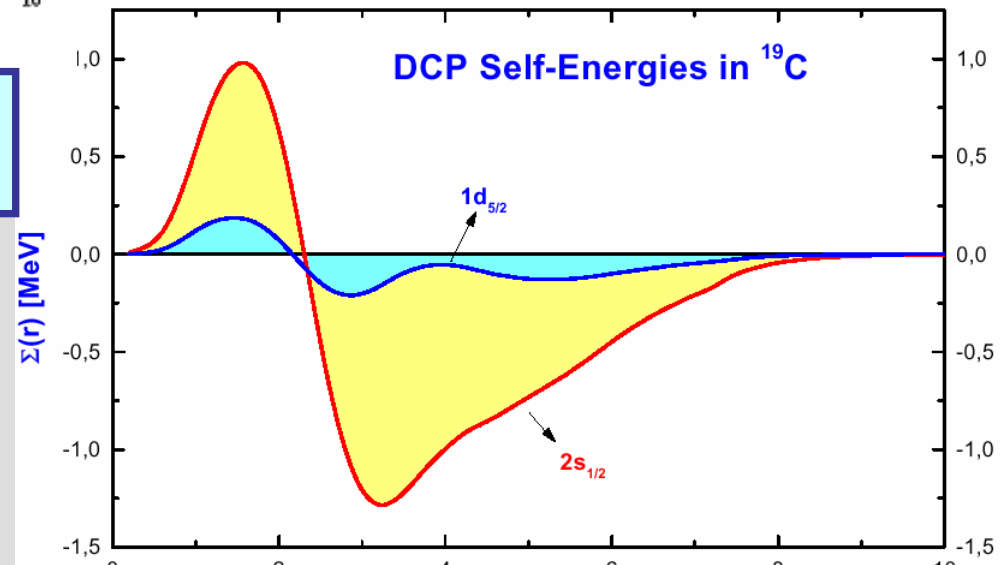
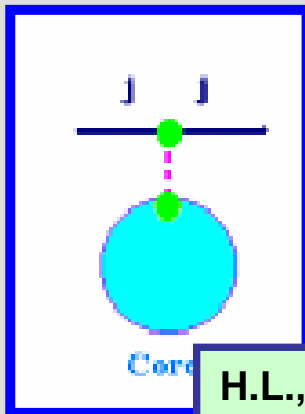
$$S_n(\lambda) = \sum_c |M_c(\lambda)|^2 E_c^n$$

The Neutron Halo in ^{19}C : Transition from Mean-Field to Correlation Dynamics



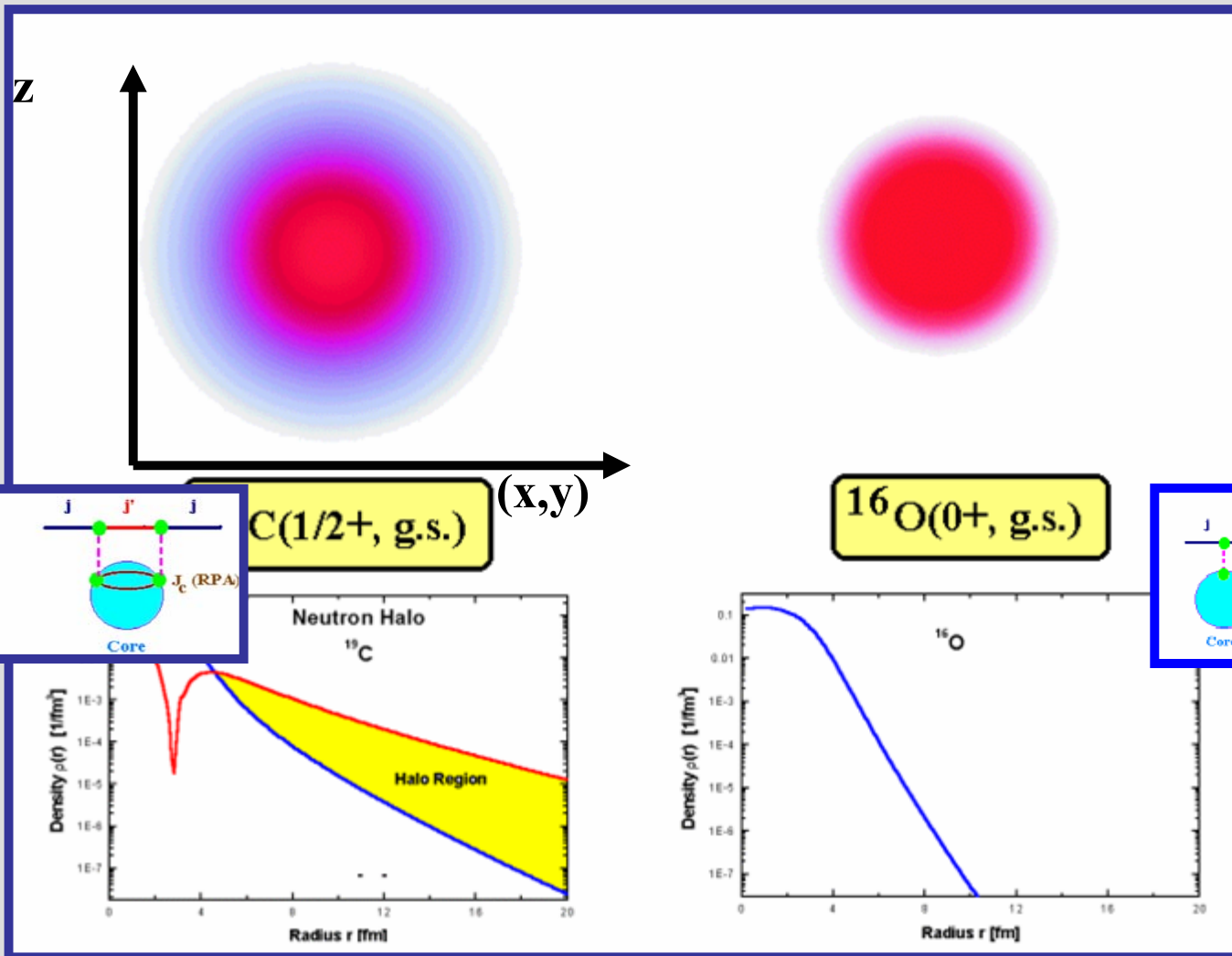
The DCP picture: Binding by Polarization Potential

The s.p. shell model picture:
Static Potential



H.L., J. Prog. Part. Nucl. 561 (2003); C. Nociforo, H.L., Nucl. Phys (2006)

Size of ^{19}C and ^{16}O



$^{19}\text{C}(1/2^+, \text{g.s.})$

$^{16}\text{O}(0^+, \text{g.s.})$

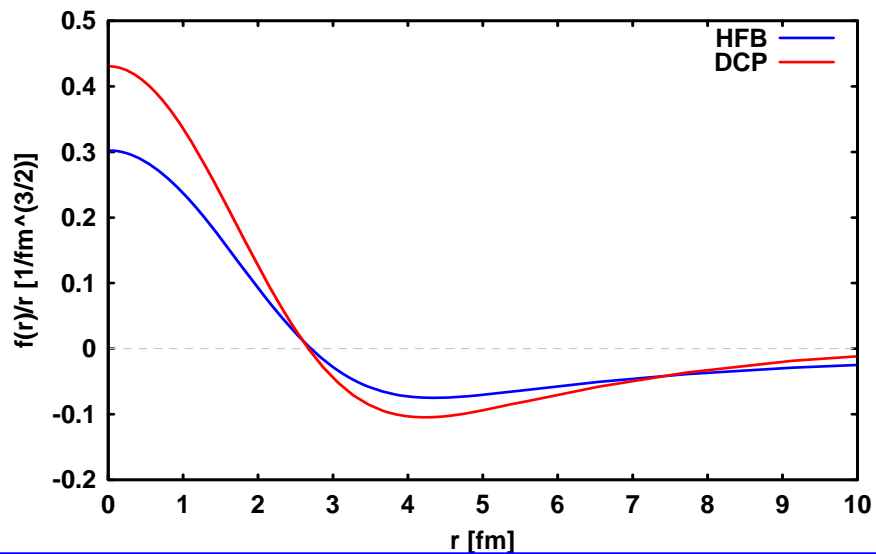
$$\sqrt{\langle r^2 \rangle (1/2^+)} = 5.34 \text{ fm}$$

$$S = 0.40$$

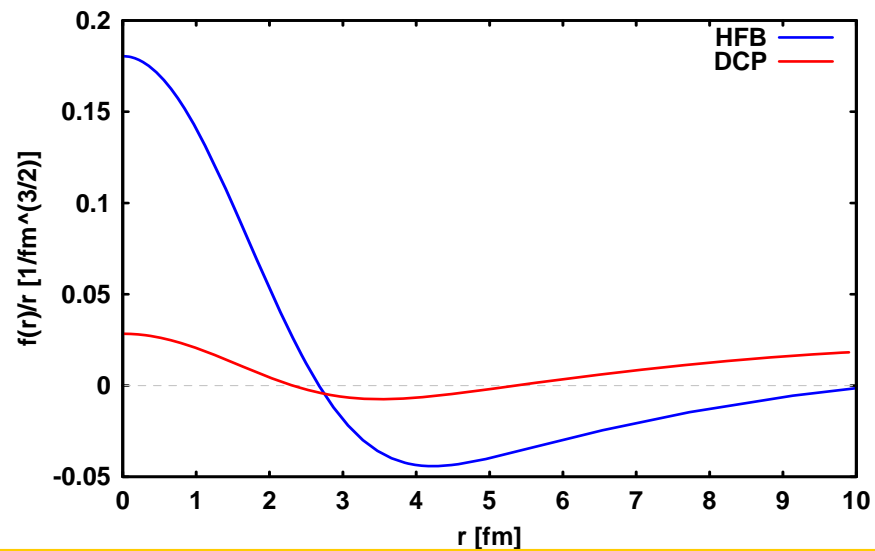
$$\sqrt{\langle r^2 \rangle (1/2^-)} = 2.96 \text{ fm}$$

$$S = 0.97$$

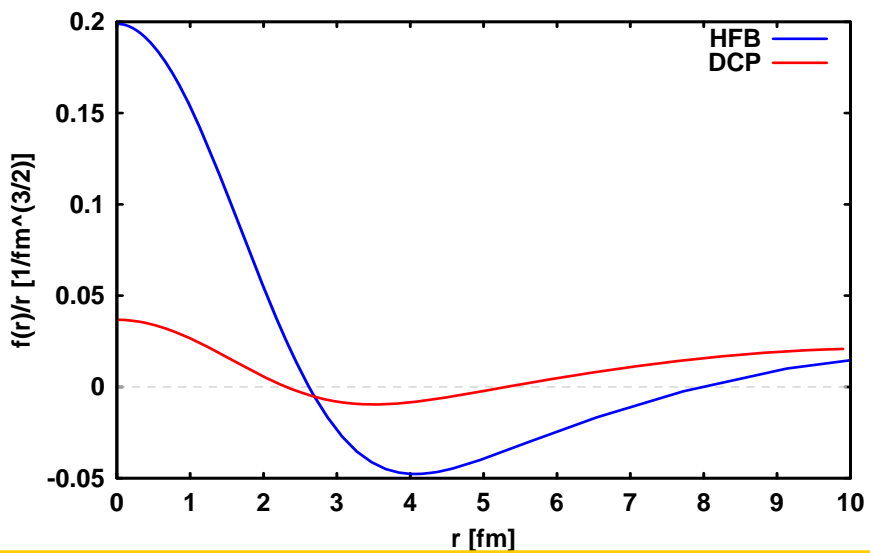
19C(1/2+) DCP Wavefunctions



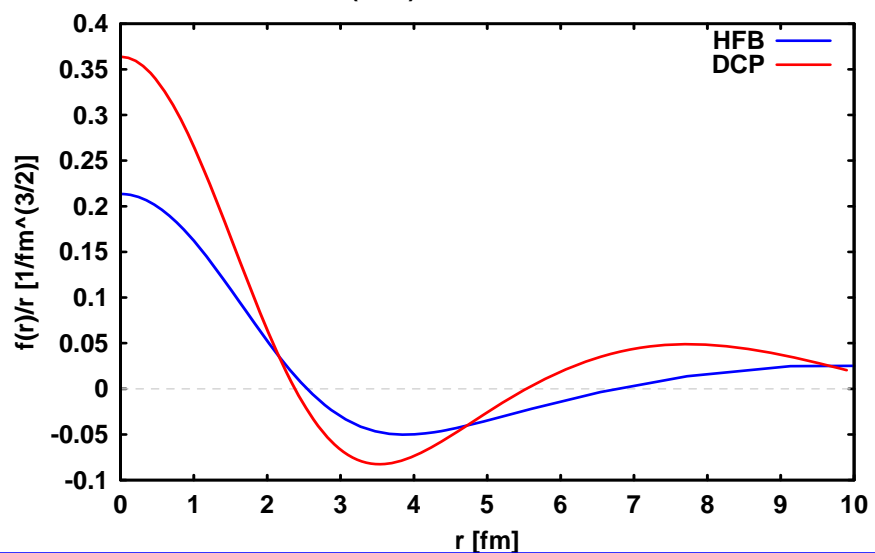
19C(1/2+) DCP Wavefunctions



19C(1/2+) DCP Wavefunctions

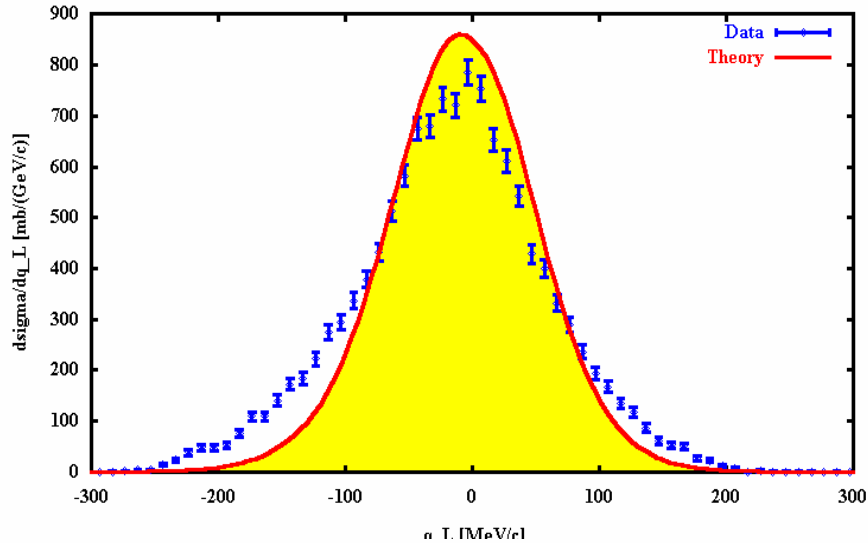


19C(1/2+) DCP Wavefunctions

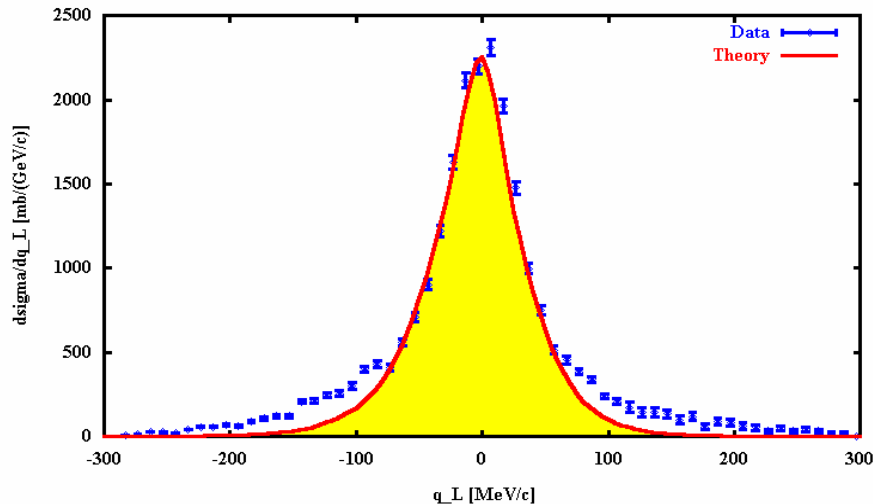


Momentum Distributions for $^{17,19}\text{C} \rightarrow ^{16,18}\text{C} + n$ Carbon Target, $E_{\text{lab}} \approx 900 \text{ A MeV}$

$^{17}\text{C} \rightarrow ^{16}\text{C}$ (on C target, 904 MeV)
 $\sigma_{\text{total}}(-1n) = 124 \text{ mb}$



$^{19}\text{C} \rightarrow ^{18}\text{C}$ (on C target, 910 MeV)
 $\sigma_{\text{total}}(-1n) = 192 \text{ mb}$



• Binding: Correlation Dynamics

• $^{17}\text{C}(5/2^+, \text{g.s.})$

• $S_n(\text{the.}) = 715 \text{ keV}$

• $C^2S(\text{g.s.}) = 0.41$

$\Gamma_{\text{theory}} : 132 \text{ MeV/c}$

$\Gamma_{\text{experim.}} : 143 \pm 5 \text{ MeV/c}$

$\sigma(-1p, \text{the.}) : 124 \text{ mb}$

$\sigma(-1p, \text{exp.}) : 129 \pm 22 \text{ mb}$

• Binding: Correlation Dynamics

• $^{19}\text{C}(1/2^+, \text{g.s.})$

• $S_n(\text{the.}) = 263 \text{ keV}$

• $C^2S(\text{g.s.}) = 0.40$

$\Gamma(\text{the.}) : 69 \text{ MeV/c}$

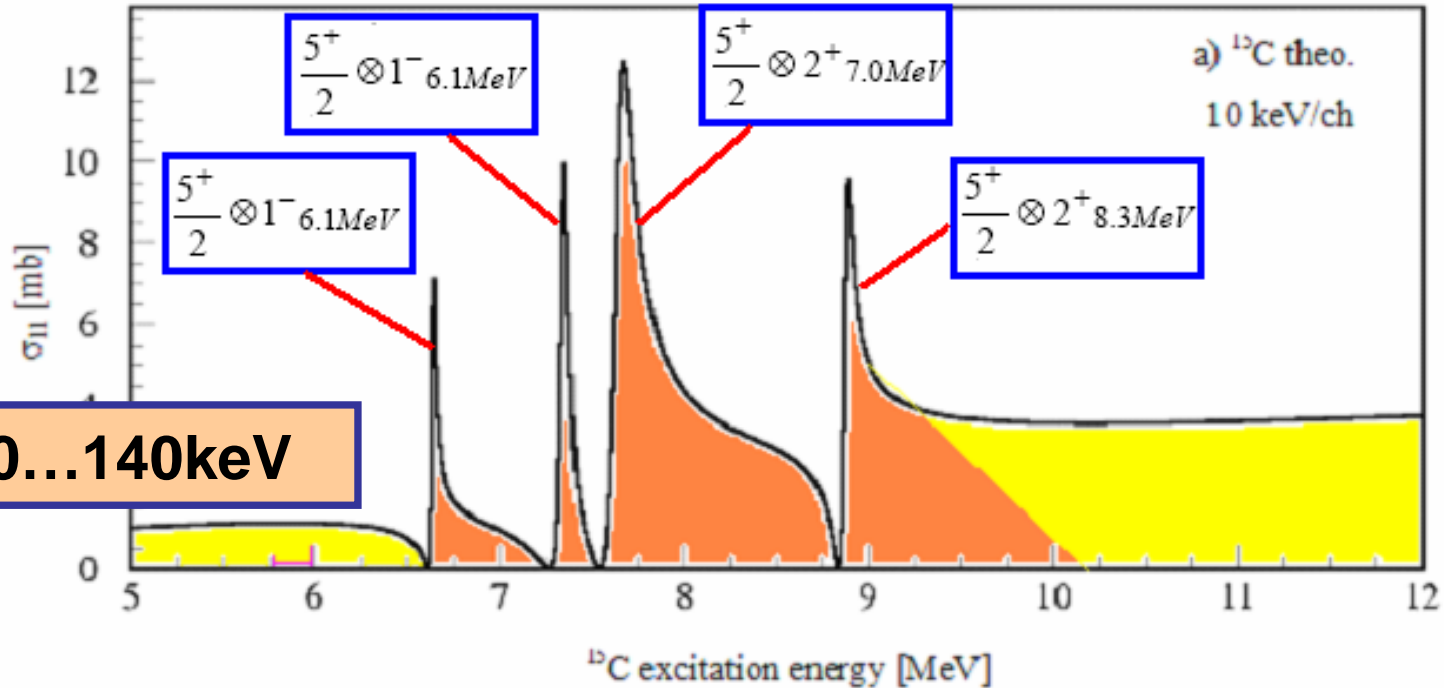
$\Gamma(\text{exp.}) : 68 \pm 3 \text{ MeV/c}$

$\sigma(-1p, \text{the.}) : 192 \text{ mb}$

$\sigma(-1p, \text{exp.}) : 233 \pm 51 \text{ mb}$

Core-excited Fano-Resonances in ^{15}C

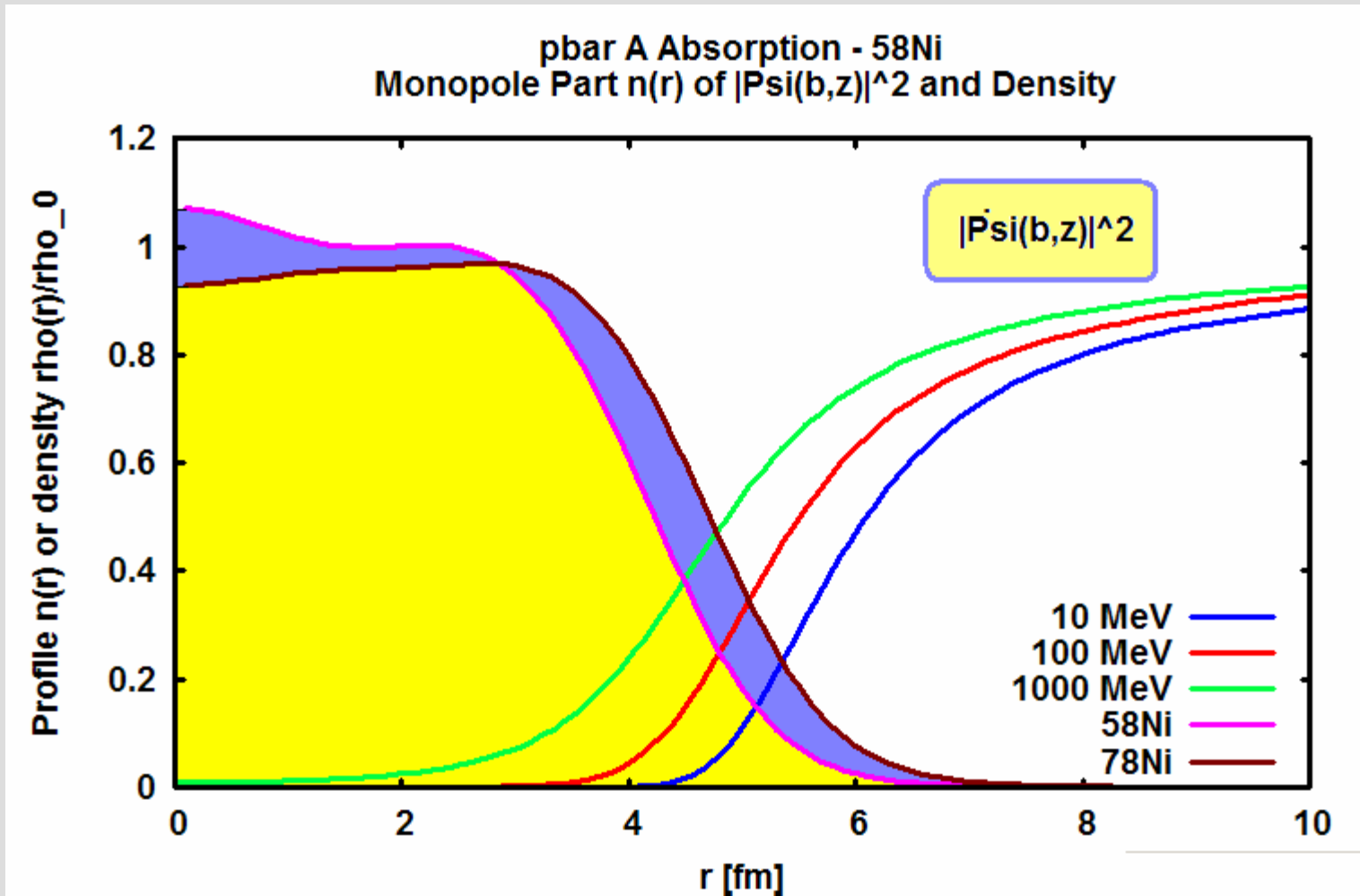
Core excitations in Continuum States:
Open Quantum System



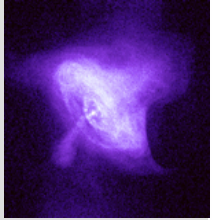
$\Gamma \sim 60 \dots 140 \text{keV}$

Sonja Orrigo, H.L. et al., Phys.Lett. B633 (2006)

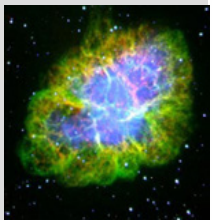
Antiproton+A Absorption: Overlap of the Reaction Kernel and the Nuclear Density



DDRH Neutron Star Mass-Radius Relation (TOV Eq.):

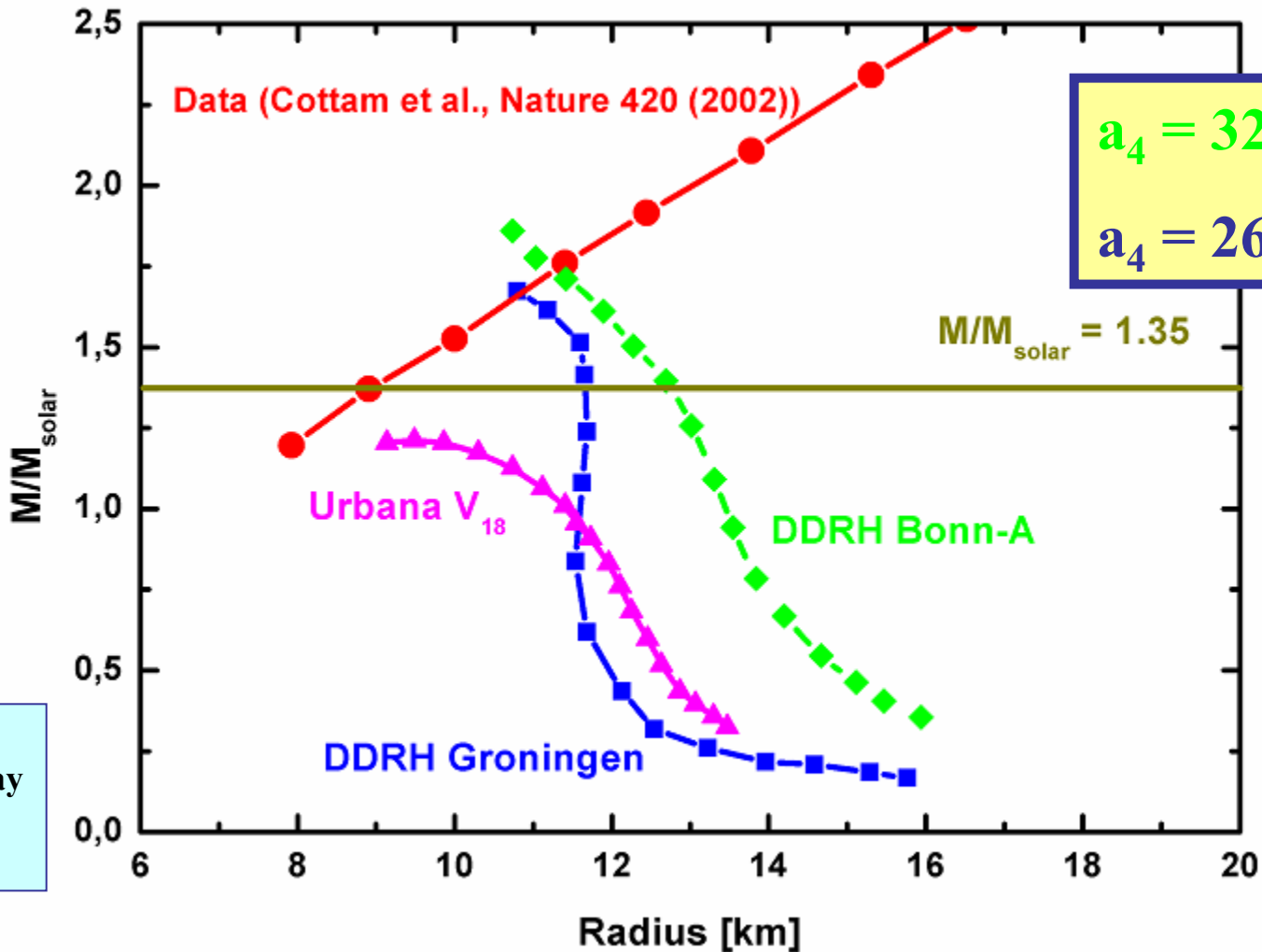


X-ray



Optical

Crab Nebula
Chandra X-Ray
Observatory
and HST



PRC 64
(2001)
025804;
H.L.
Springer
Lect.
Notes
(2004)

Data from the XMM-Newton X-ray space observatory:

Gravitational Red-Shift $z \sim M/R$

(Fe-Lines from a series of 28 X-ray bursts from EXO07481676)

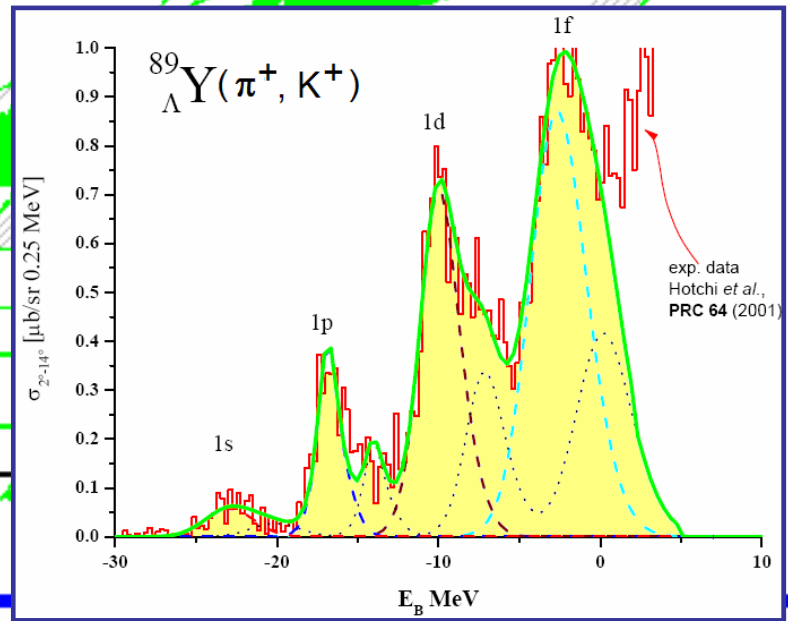
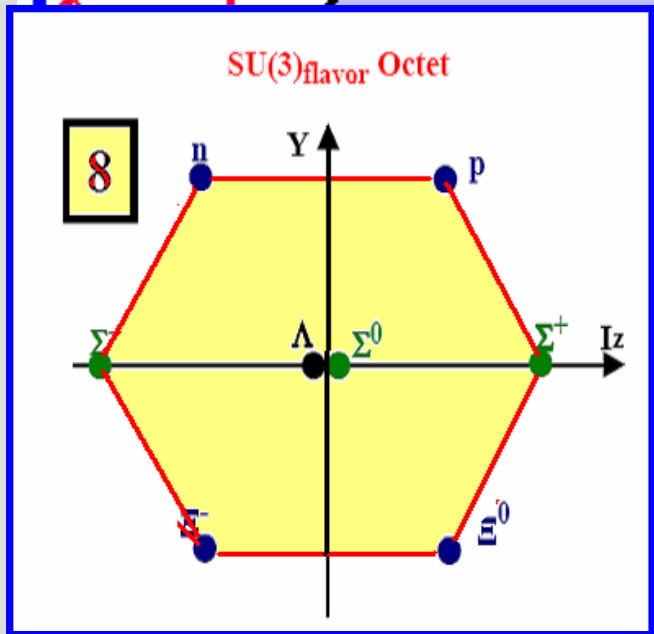
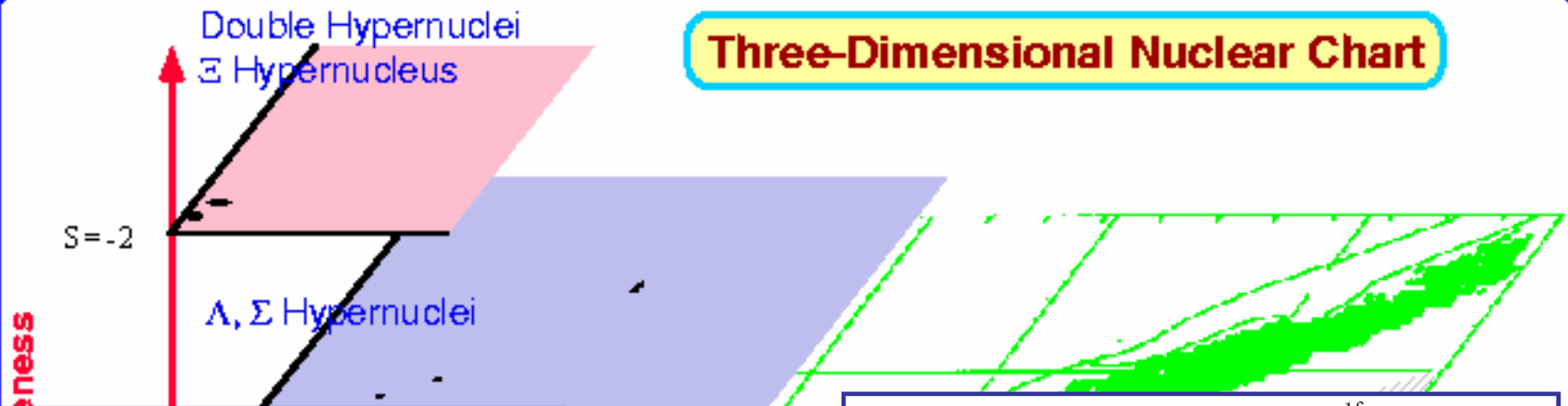
Summary and Outlook:

- *ab initio* Description of Stable and Unstable Nuclei
- Interactions in Asymmetric Nuclear Matter
- Continuum Effects: Open Quantum Systems
- Softness, Polarizability of Exotic Nuclei and LRC
- Mean-field Dynamics or Correlation Dynamics?
- Relation to Astrophysics: Neutron Stars
- Unique: Nucleo-Engineering on the Femto-scale!

Thanks to all participants from the organizers Tom, Roy, H., and Jeff for the lively discussions and the productive engagement

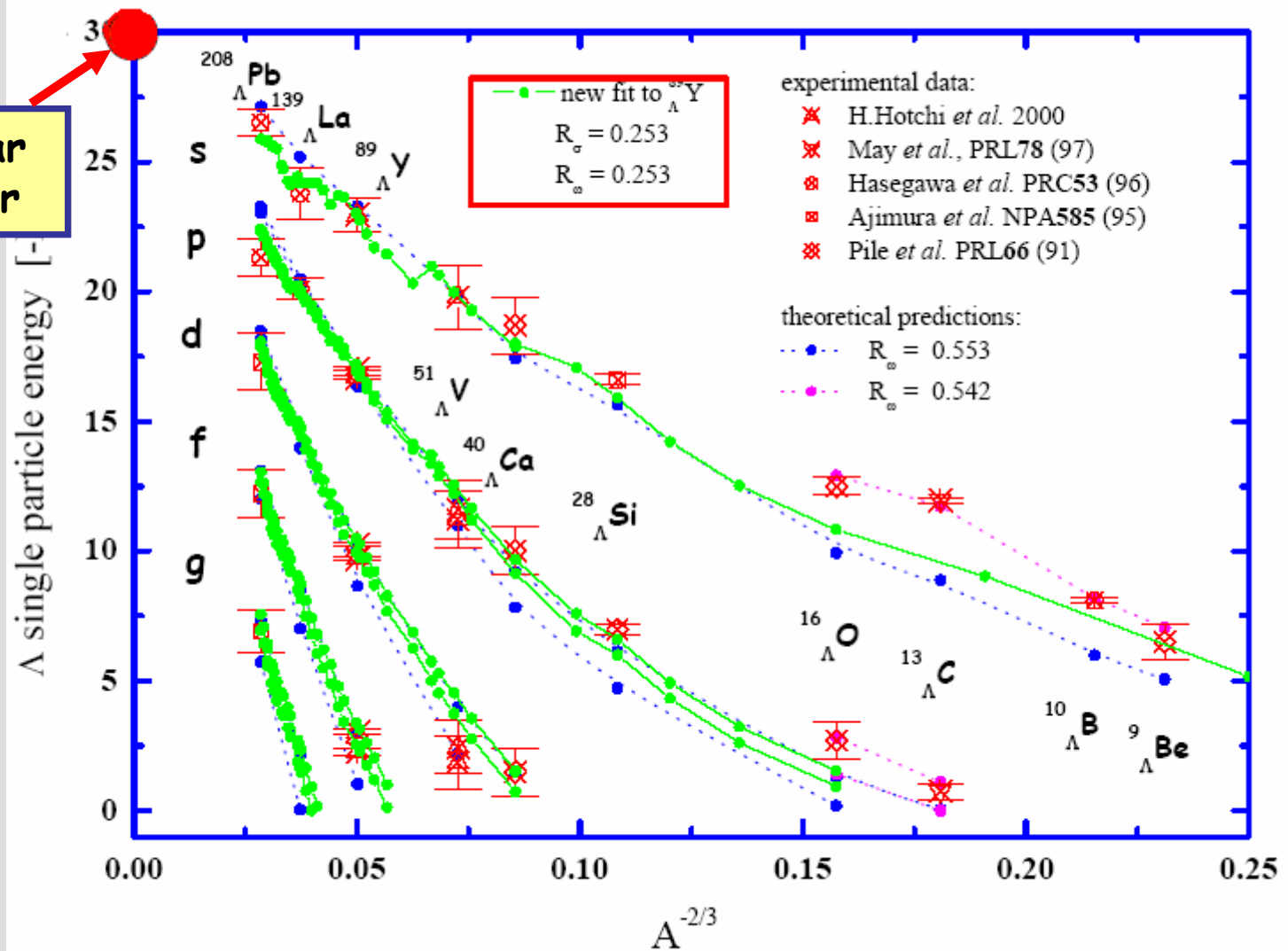
Strangeness and Hypernuclear Physics: From $SU(2)$ Isospin to $SU(3)$ Flavour Dynamics

Three-Dimensional Nuclear Chart



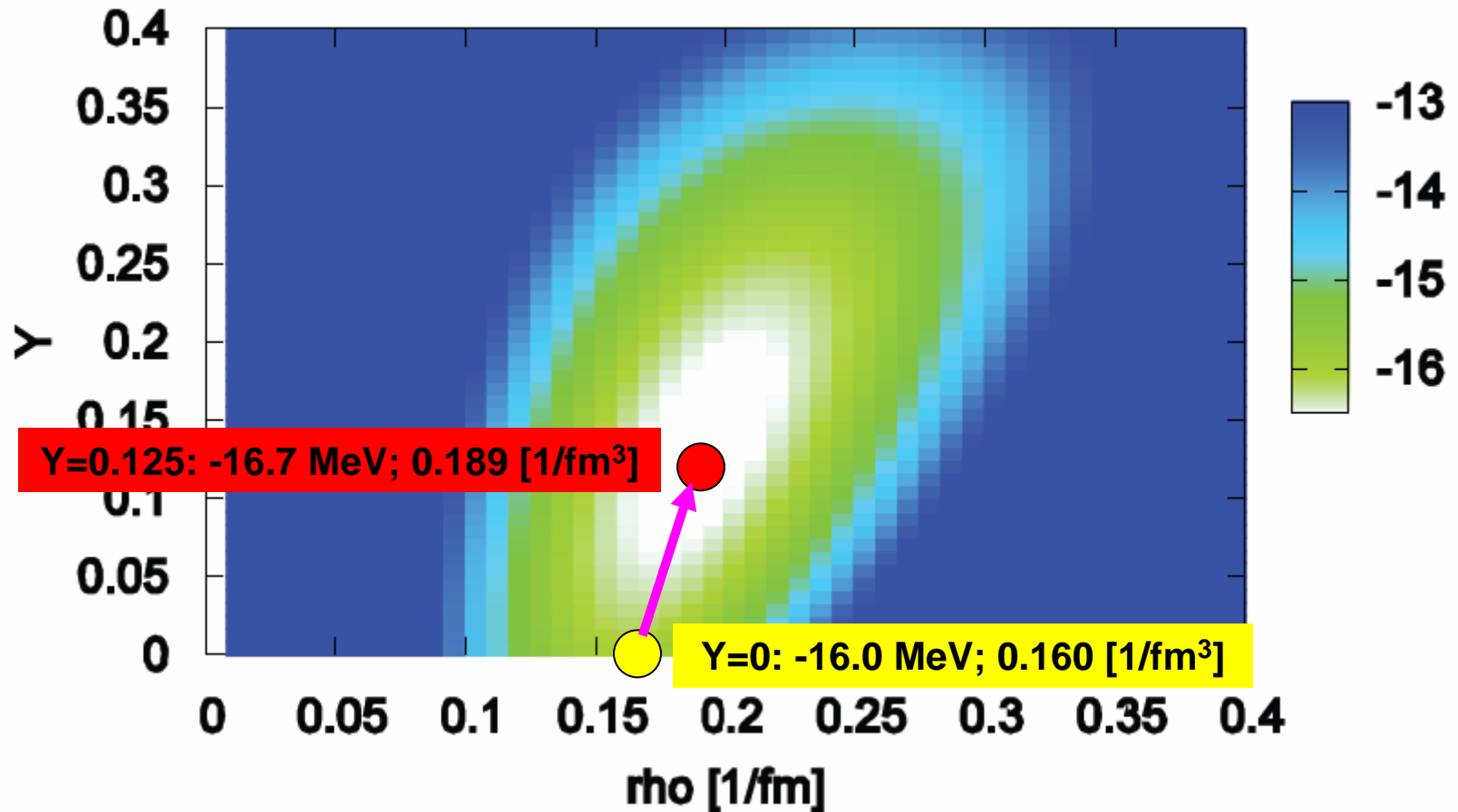
DDRH Flavour Dynamics: Λ Single Particle Energies

Nuclear Matter



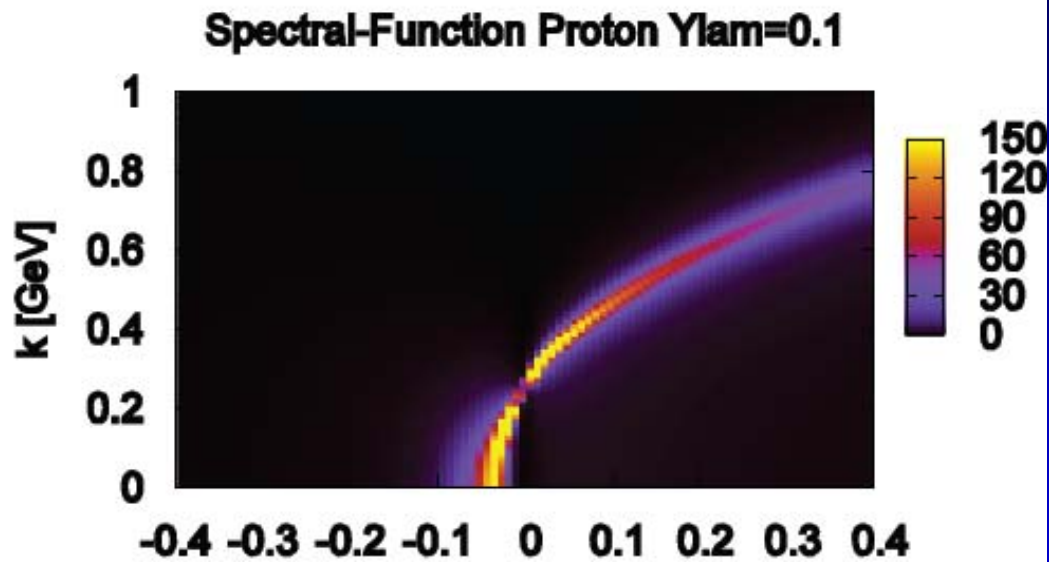
DDRH Theory: Density Dependent NN and $N\Lambda$ Dirac-Brueckner Vertices

Binding Energy of Infinite Hypermatter

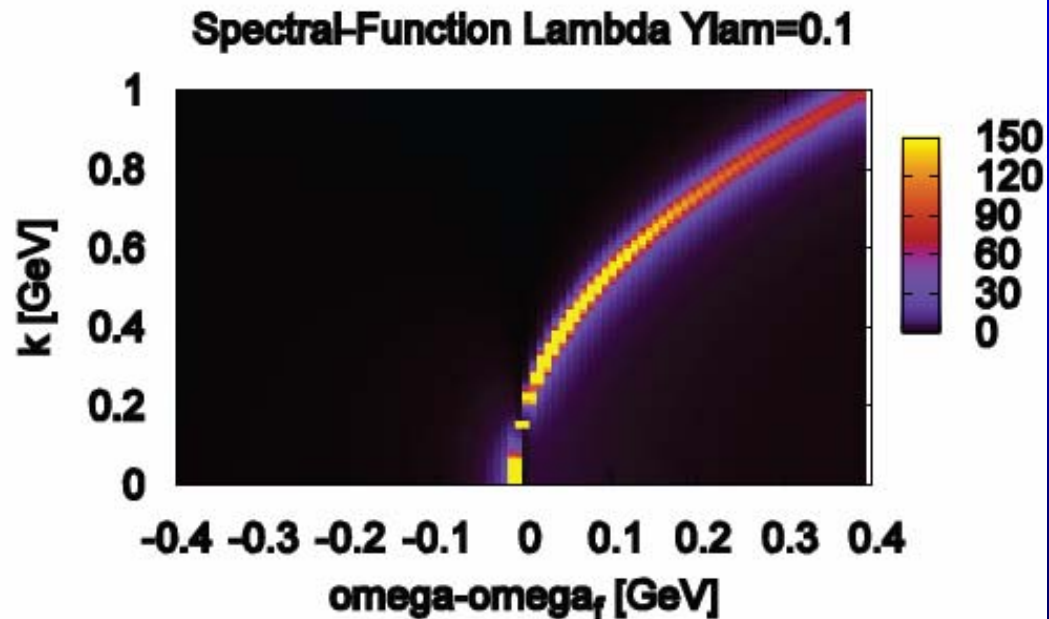


$$B(\rho, Y) = E(\rho, Y) / \rho - YM_{\Lambda} - (1 - Y)M_N$$

$$S_p(k, \omega)$$

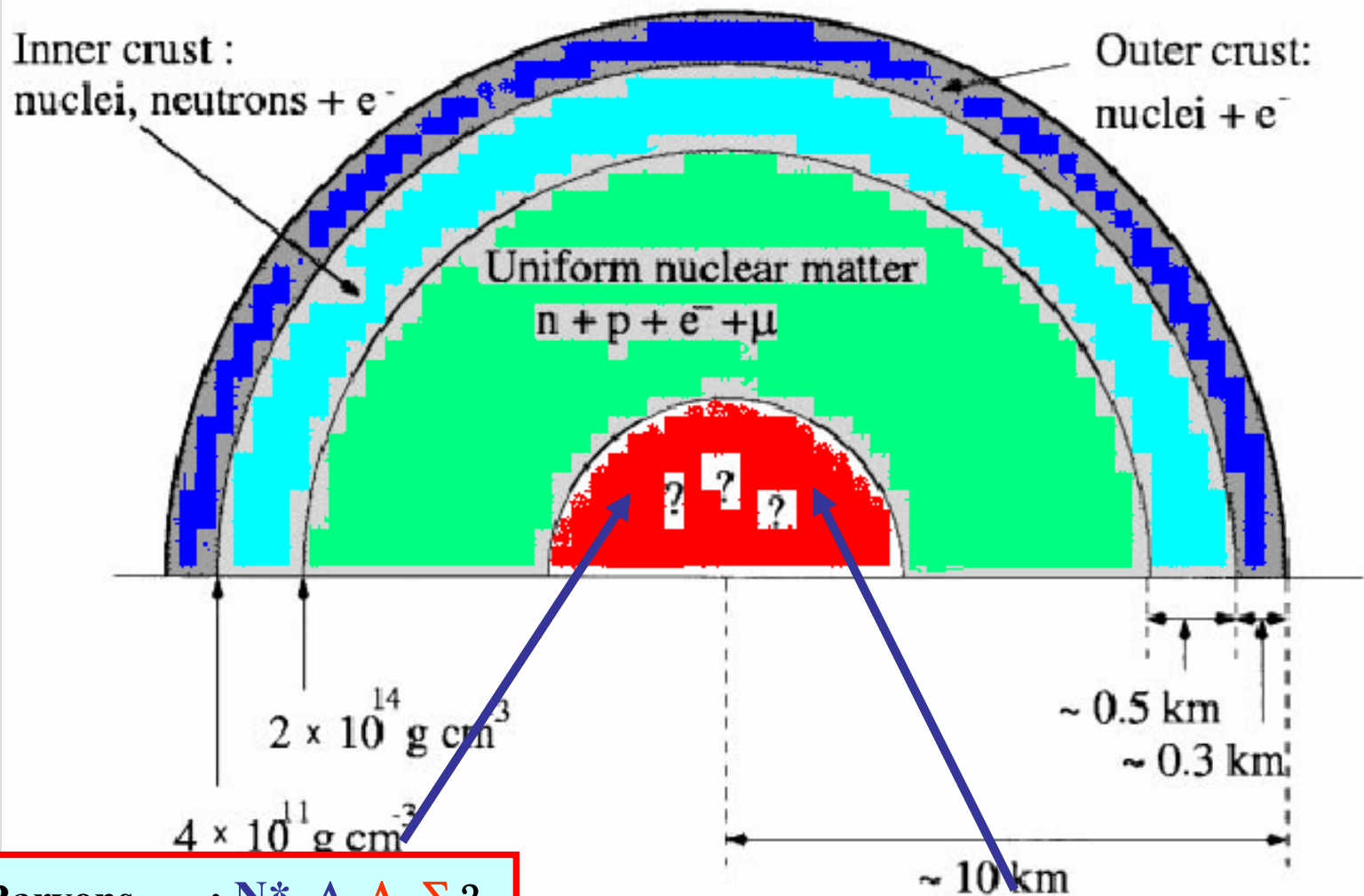


$$S_{\Lambda}(k, \omega)$$



Beyond
Mean-Field
Dynamics:
Spectral
Functions in
Infinite
Hypermatter

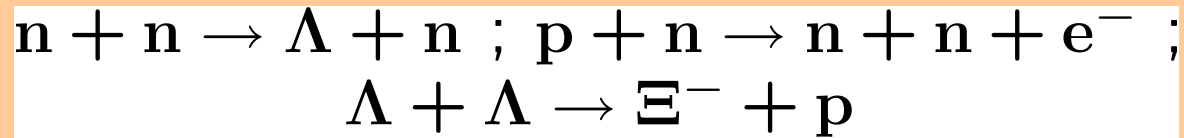
Expected Structure of a Neutron Star



Baryons : N^* , Δ , Λ , Σ ?
Condensates: π , K ... ?

Quark Matter, QGP, CFL?

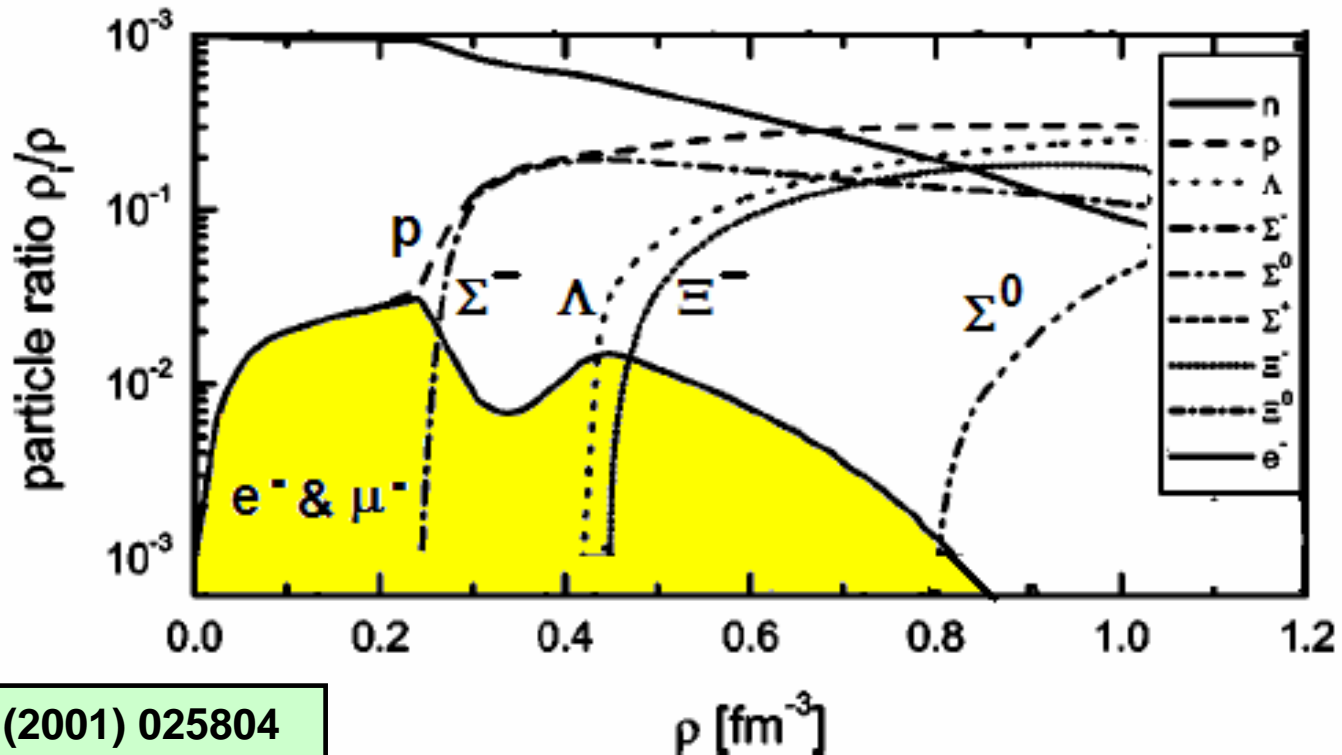
Charge-Neutral Neutron Star Matter in β -Equilibrium



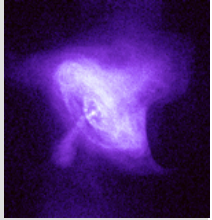
$$\mu_i = b_i \mu_n - q_i \mu_e$$

Creation of Strangeness:

$\rho \sim 2\rho_0$: hyperon threshold (Σ^{-} , Λ), $\rho > 5\rho_0$: hypermatter dominates



DDRH Neutron Star Mass-Radius Relation (TOV Eq.):

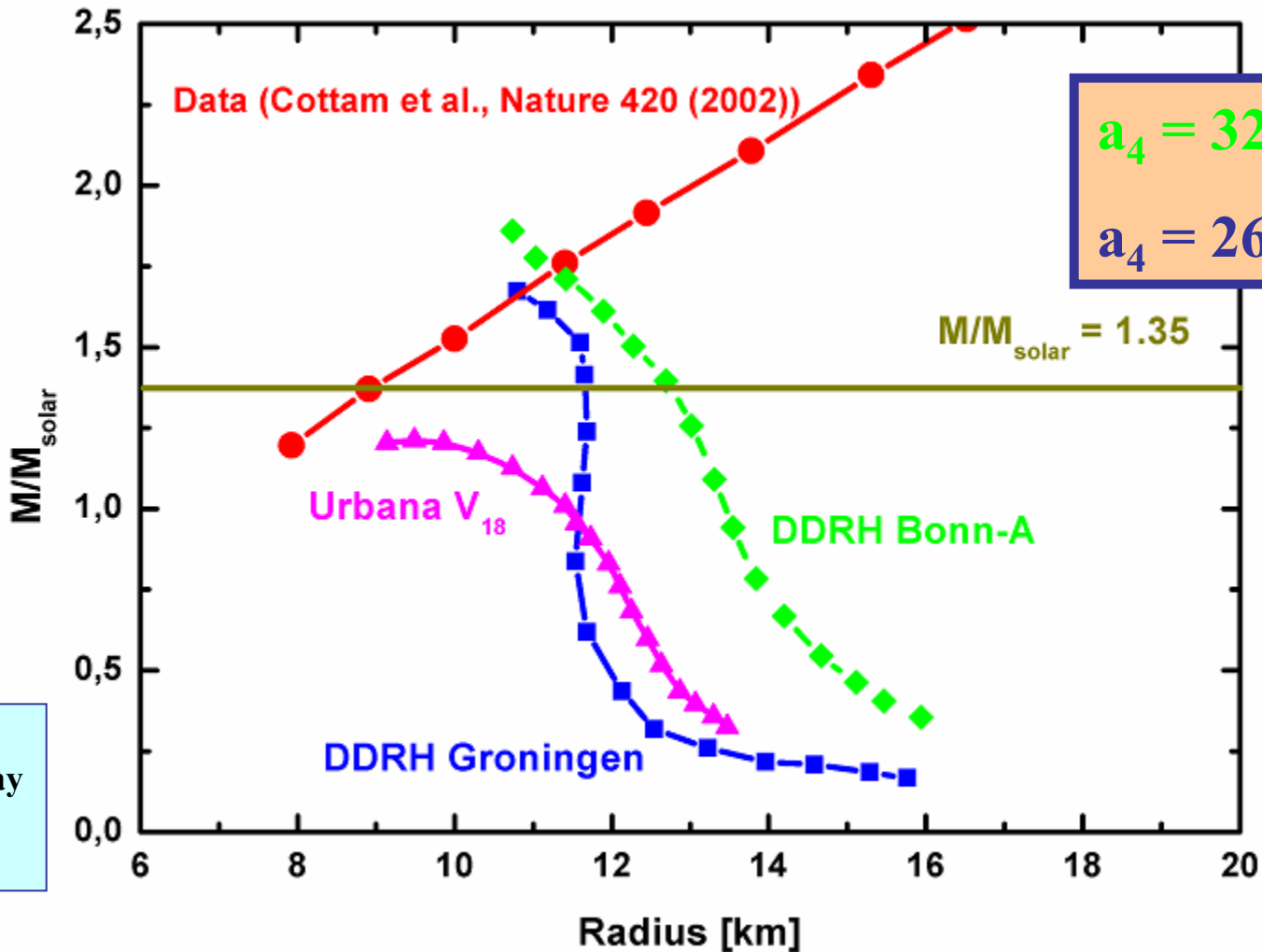


X-ray



Optical

Crab Nebula
Chandra X-Ray
Observatory
and HST



$a_4 = 32 \text{ MeV}$

$a_4 = 26 \text{ MeV}$

PRC 64
(2001)
025804;
H.L.
Springer
Lect.
Notes
(2004)

Data from the XMM-Newton X-ray space observatory:

Gravitational Red-Shift $z \sim M/R$

(Fe-Lines from a series of 28 X-ray bursts from EXO07481676)